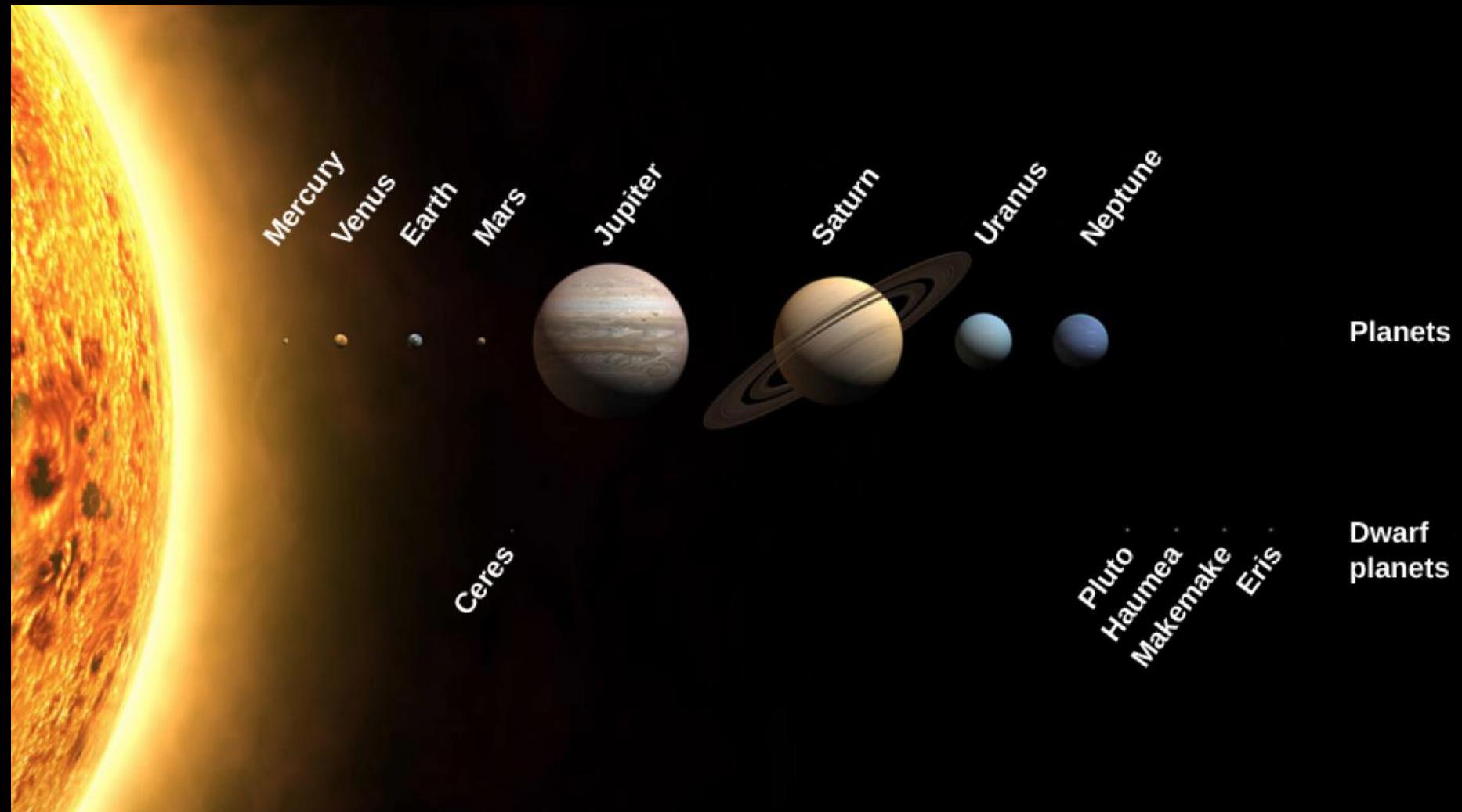


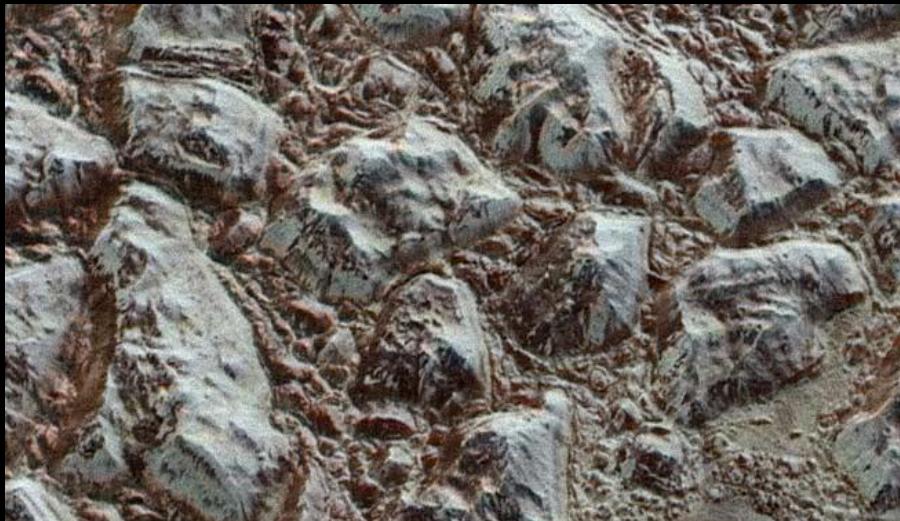
Our solar system!



Mars or Earth?



Pluto or Earth?

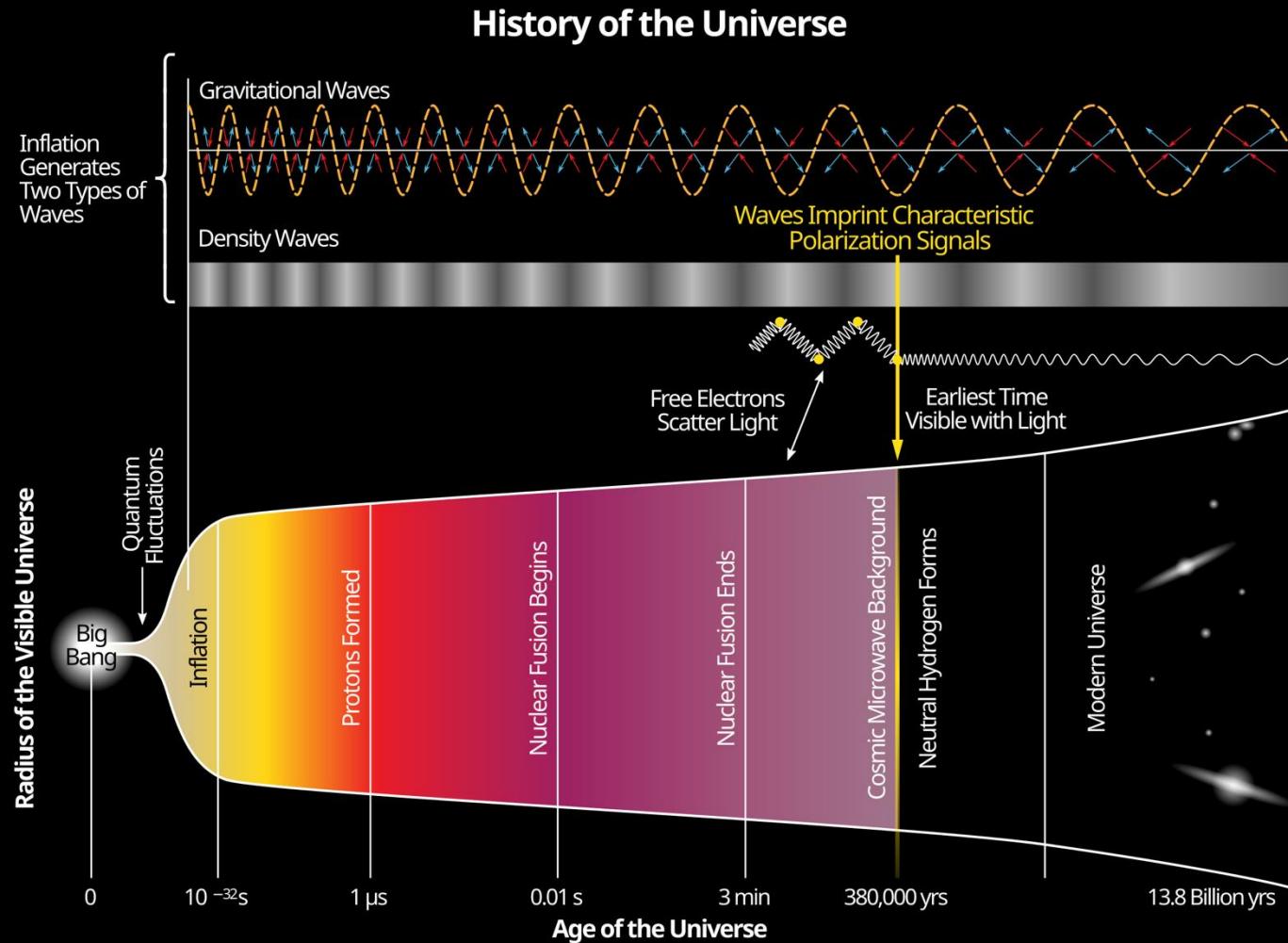




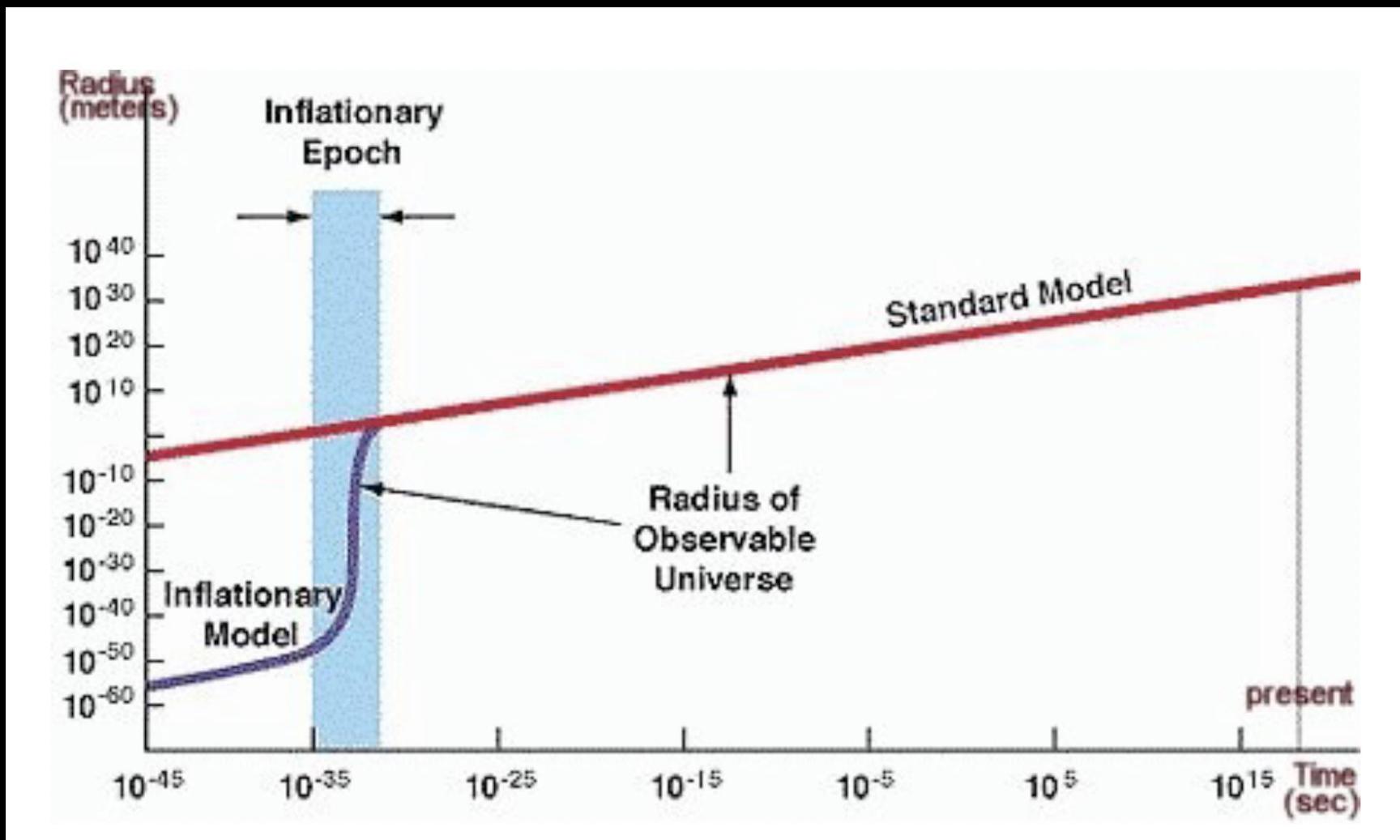
NASA/Dragonfly Titan Mission (artist's
image, planned for late 2020s)

The history of the universe (in 13 slides)

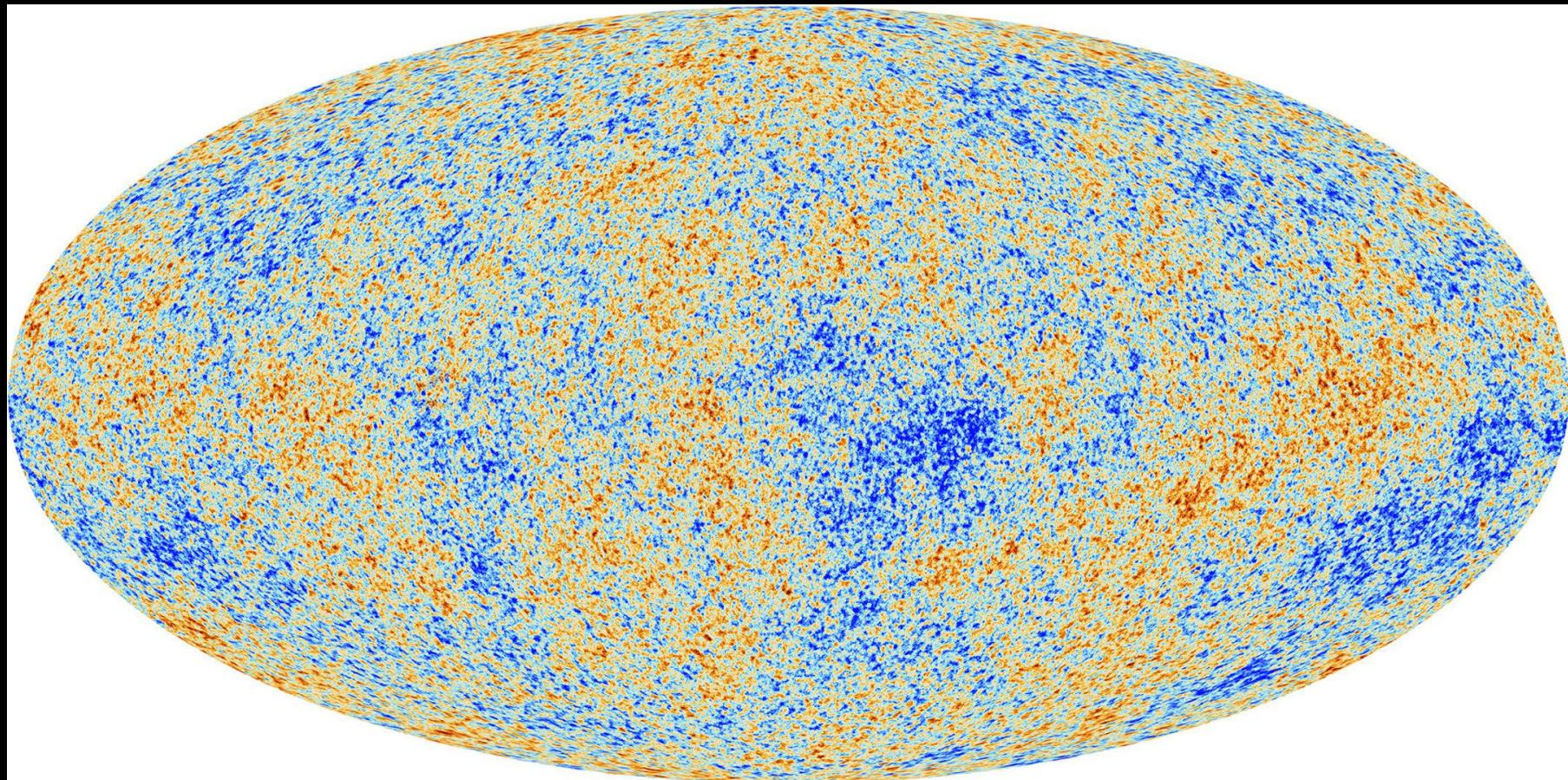
The Big Bang: an explosion of space-time!



Inflation: factor of 10^{78} from 10^{-36} to 10^{-32} s



Cosmic microwave background

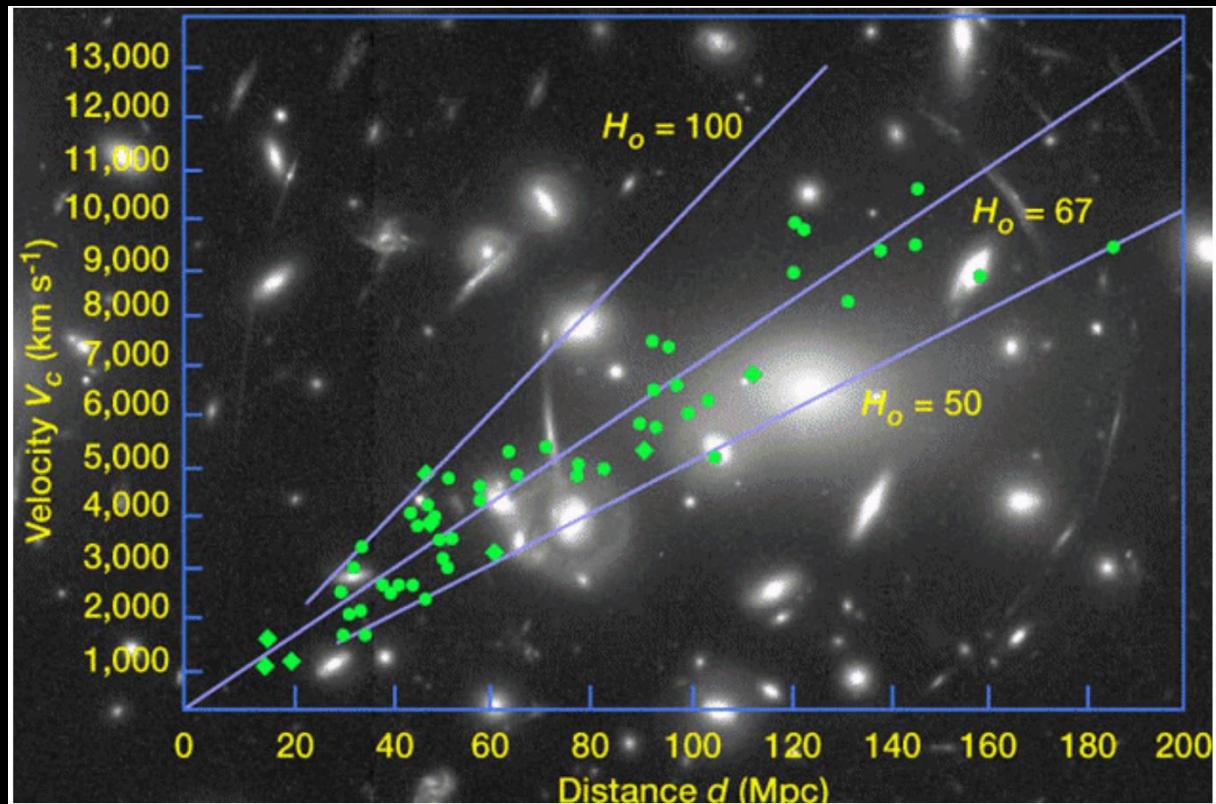


Near-uniform microwave background (smooth to 1 part in 10^4)
But not perfect: anisotropies!



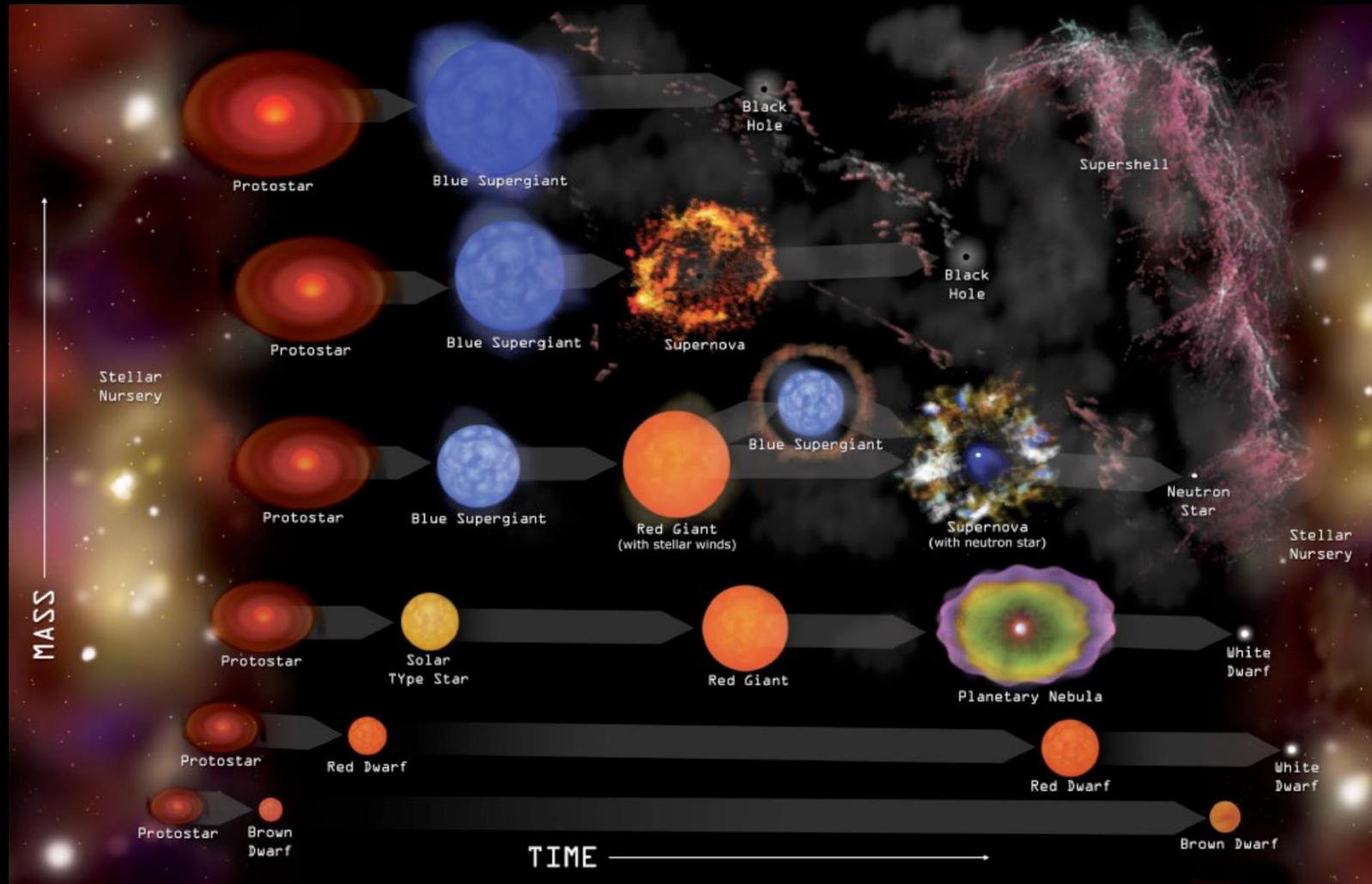
First stars and
galaxies form

Cosmic expansion and Hubble's Law



Most distant objects: redshifted (so appears red)
Distance: looks at the distance when younger (light travel time)

First stars are massive and explode!



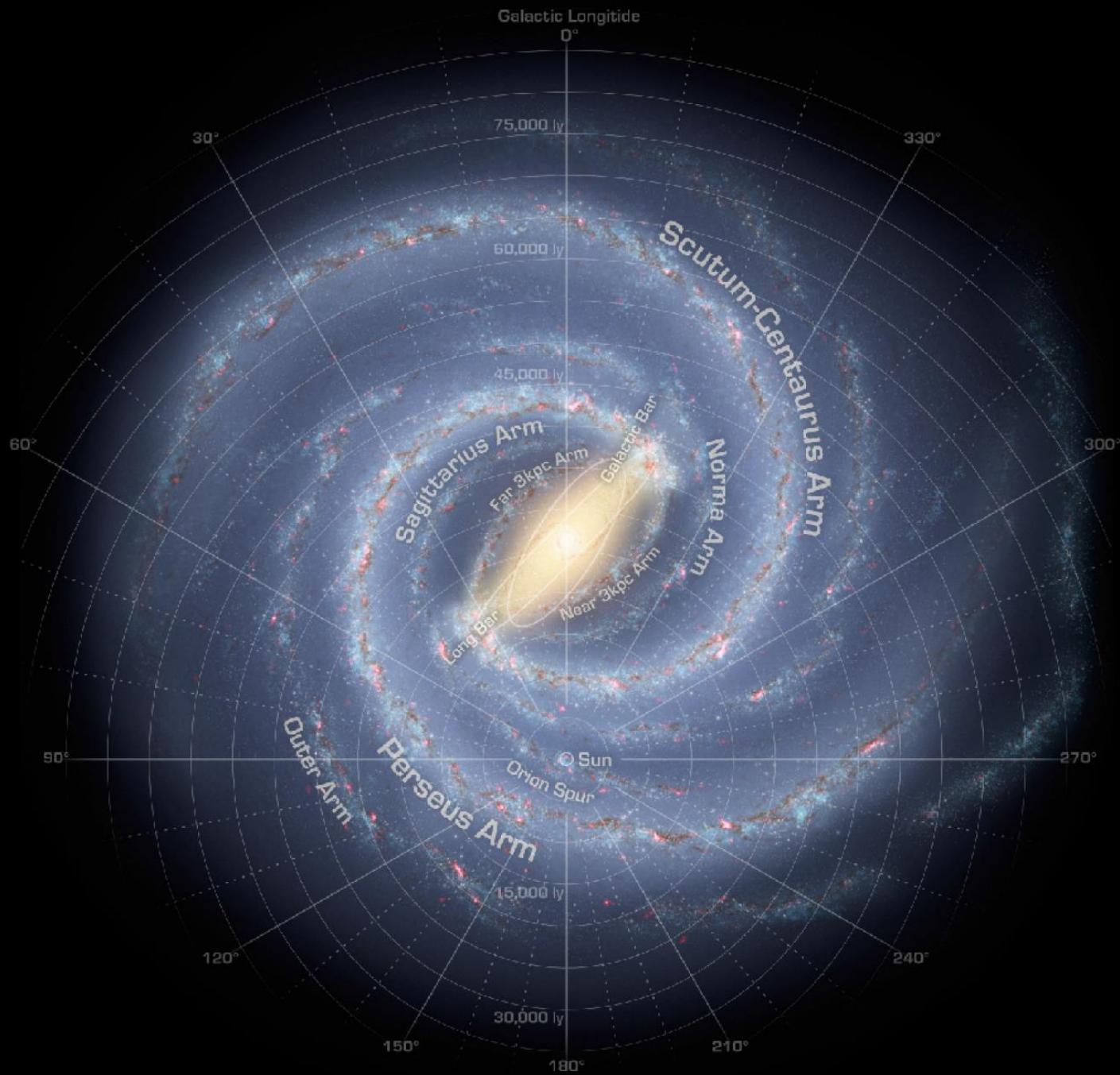
Supernova 1987A (brightest in modern times)

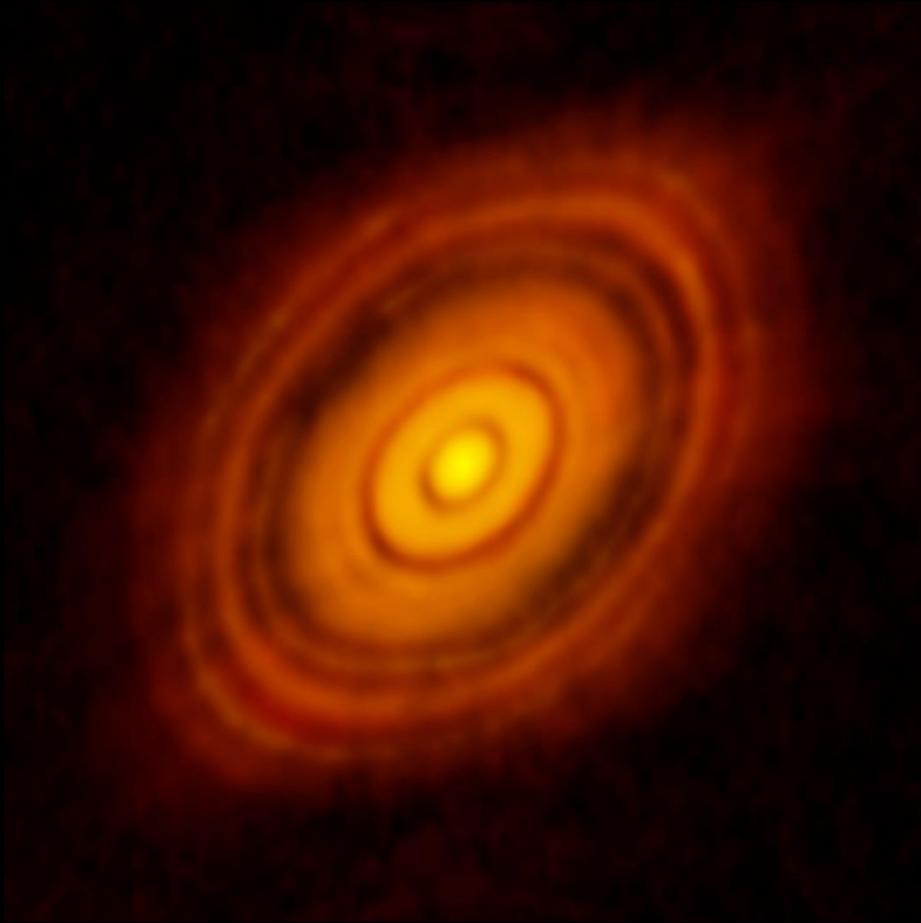


© Anglo-Australian Observatory

The Origin of the Solar System Elements

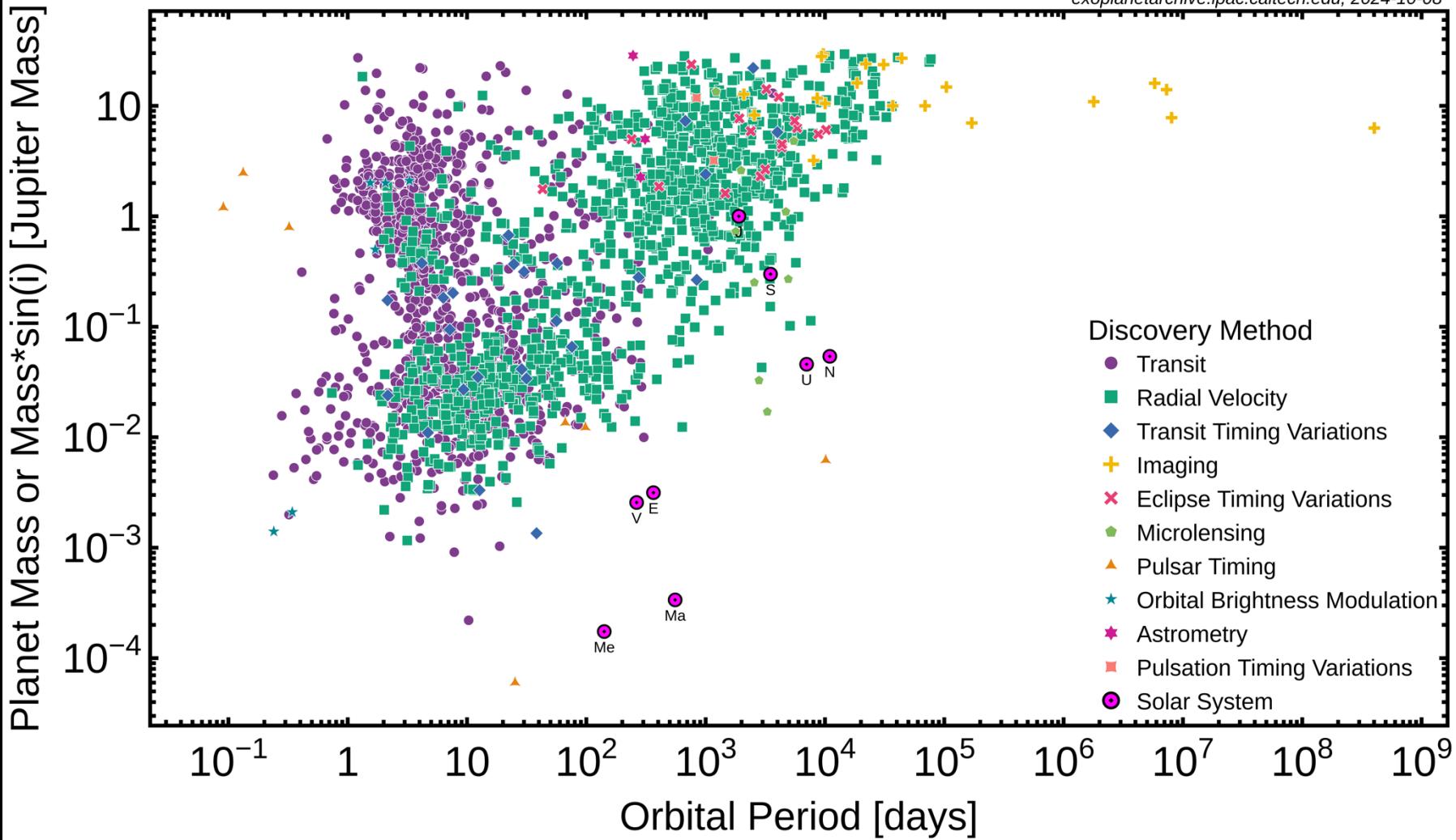
1 H	big bang fusion 					cosmic ray fission 					2 He						
3 Li	4 Be	merging neutron stars? 			exploding massive stars 			5 B	6 C	7 N	8 O	9 F	10 Ne				
11 Na	12 Mg	dying low mass stars 			exploding white dwarfs 			13 Al	14 Si	15 P	16 S	17 Cl	18 Ar				
19 K	20 Ca	21 Sc	22 Ti	23 V	24 Cr	25 Mn	26 Fe	27 Co	28 Ni	29 Cu	30 Zn	31 Ga	32 Ge	33 As	34 Se	35 Br	36 Kr
37 Rb	38 Sr	39 Y	40 Zr	41 Nb	42 Mo	43 Tc	44 Ru	45 Rh	46 Pd	47 Ag	48 Cd	49 In	50 Sn	51 Sb	52 Te	53 I	54 Xe
55 Cs	56 Ba	72 Hf	73 Ta	74 W	75 Re	76 Os	77 Ir	78 Pt	79 Au	80 Hg	81 Tl	82 Pb	83 Bi	84 Po	85 At	86 Rn	
87 Fr	88 Ra	57 La	58 Ce	59 Pr	60 Nd	61 Pm	62 Sm	63 Eu	64 Gd	65 Tb	66 Dy	67 Ho	68 Er	69 Tm	70 Yb	71 Lu	
Very radioactive isotopes; nothing left from stars																	

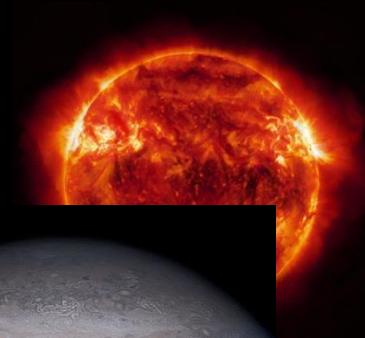
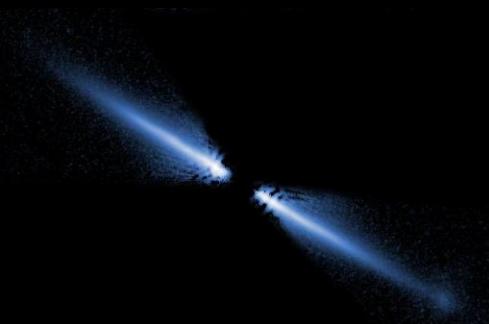




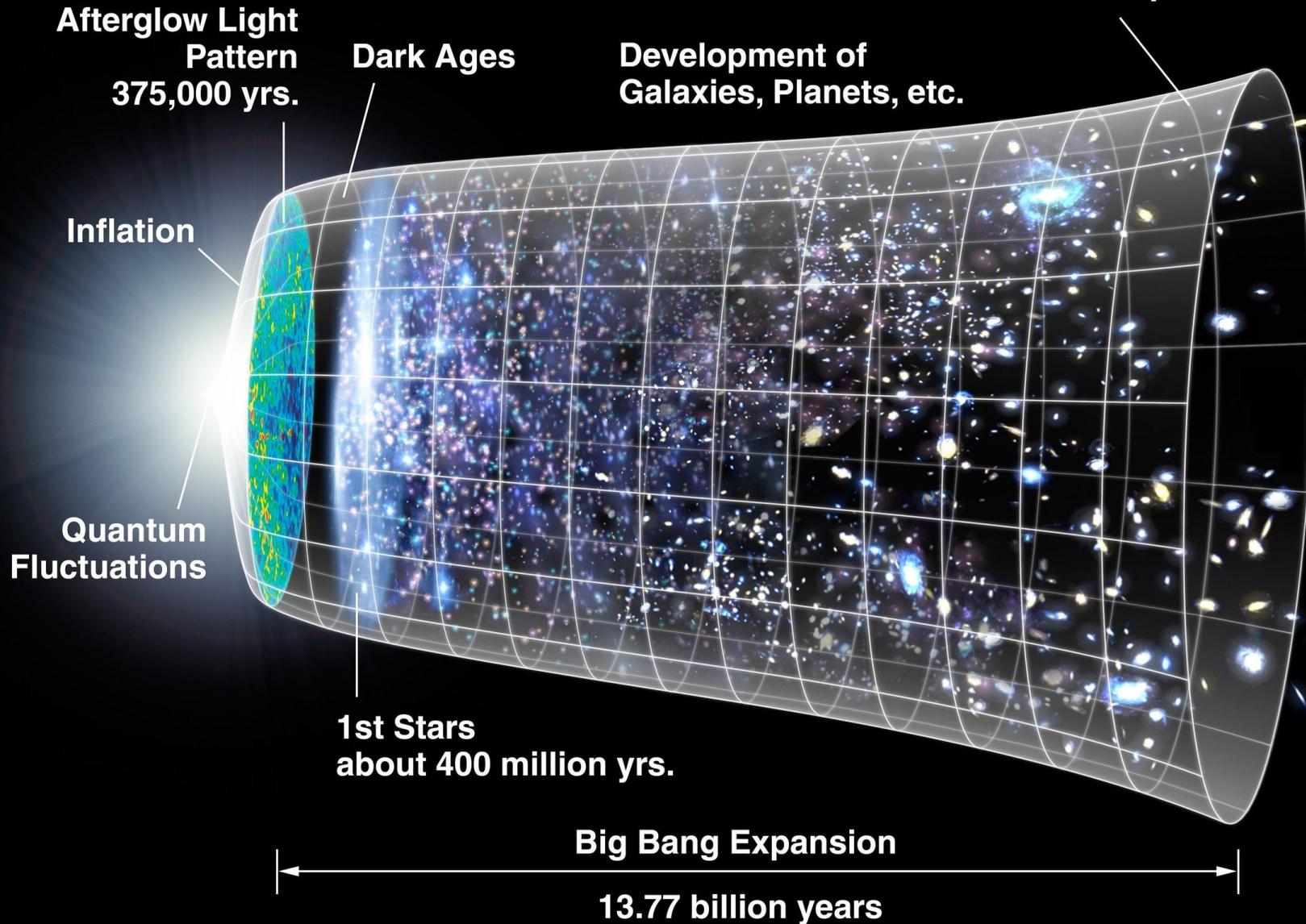
Planet Mass or Mass*sin(i) vs Orbital Period

exoplanetarchive.ipac.caltech.edu, 2024-10-08





**Dark Energy
Accelerated Expansion**





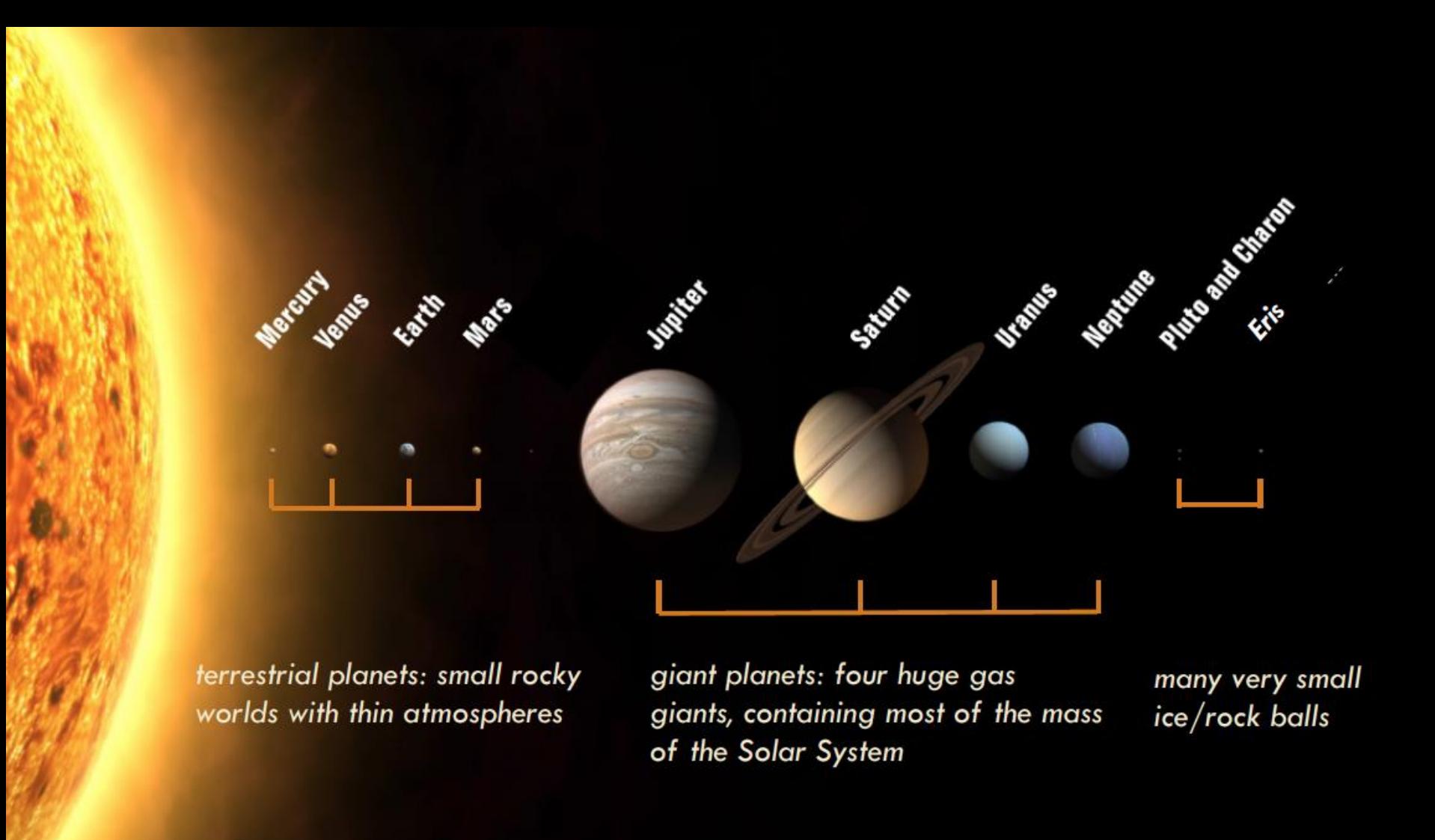
Big Bang occurs.

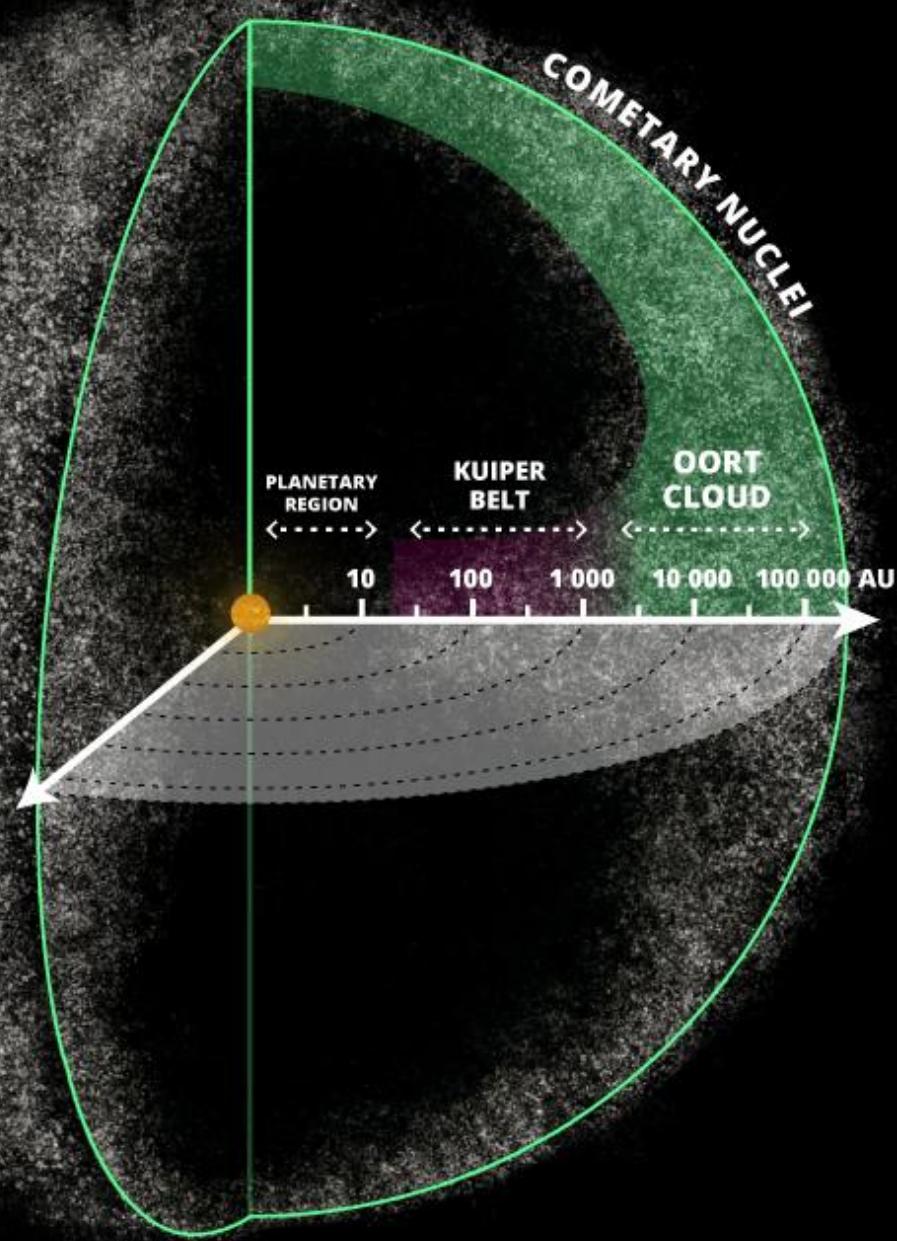
Milky Way Galaxy forms.

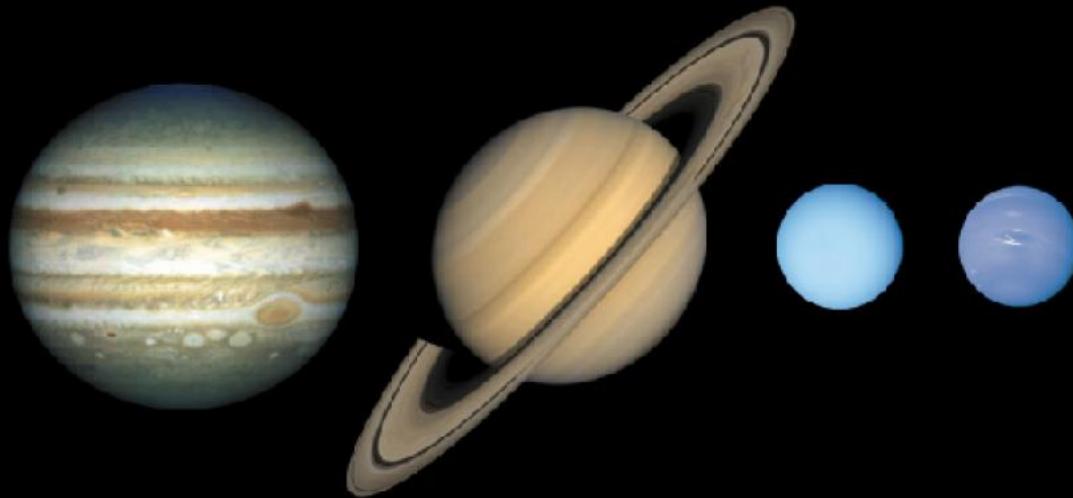
Our solar system forms. Life on Earth begins. Earth's atmosphere becomes oxygenated. First complex life forms appear.

December

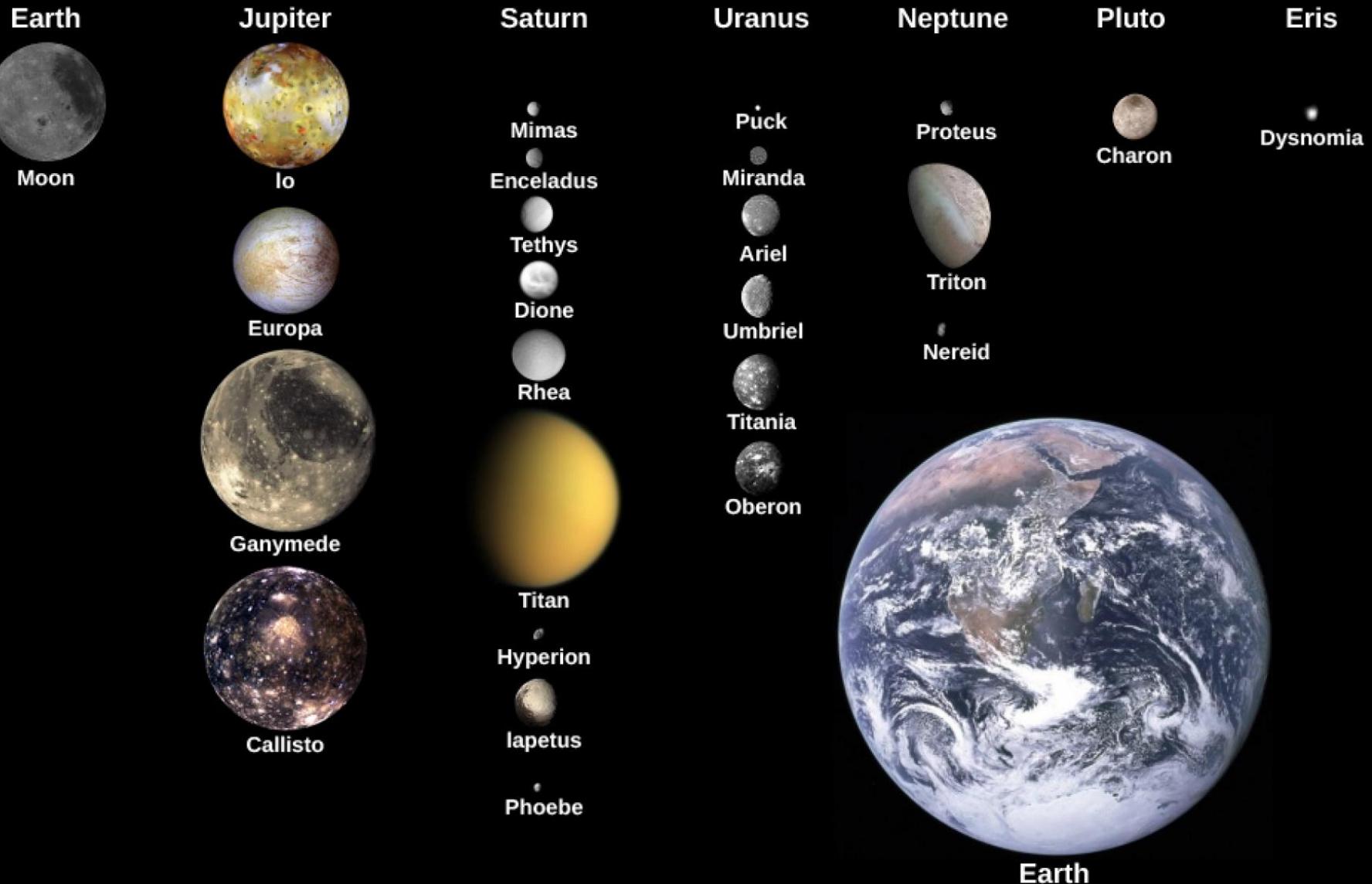
1	2	3	4	5	6	7
8	9	10	11	12	13	14
15	16	17	18	19 Vertebrates appear.	20 Land plants appear.	21
22	23	24	25 Dinosaurs appear.	26 Mammals appear.	27	28
29	30 Dinosaurs become extinct.	31 Humans appear.				







⊕ Earth





Ganymede
5262 km



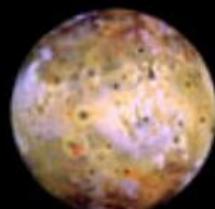
Titan
5150 km



Mercury
4880 km



Callisto
4806 km



Io
3642 km



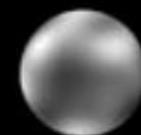
Moon
3476 km



Europa
3138 km



Triton
2706 km

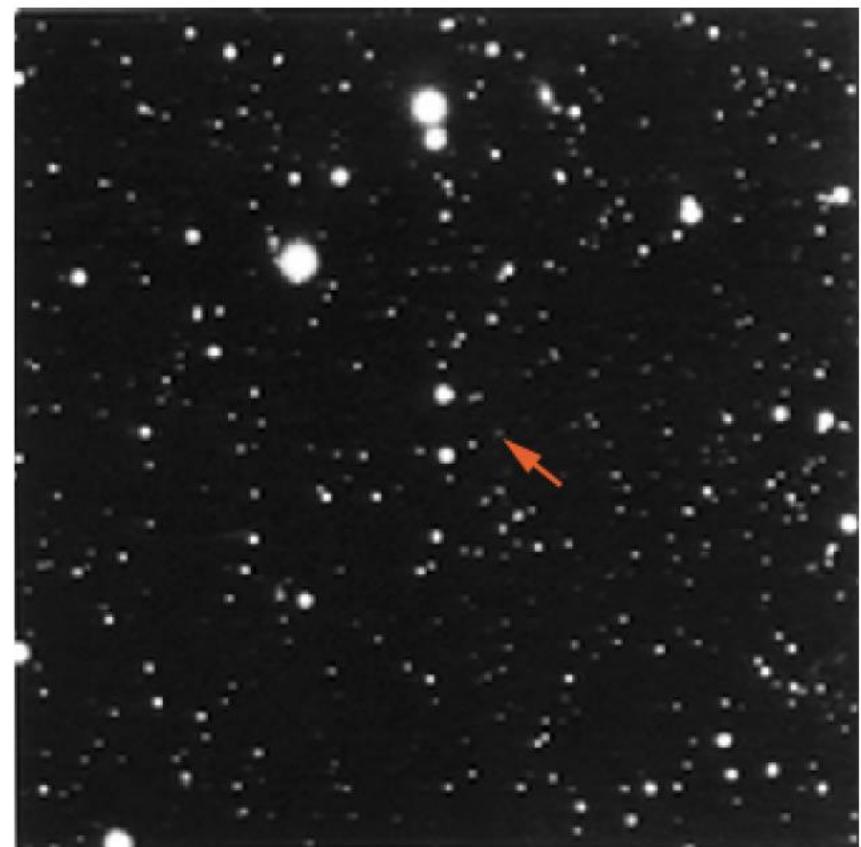
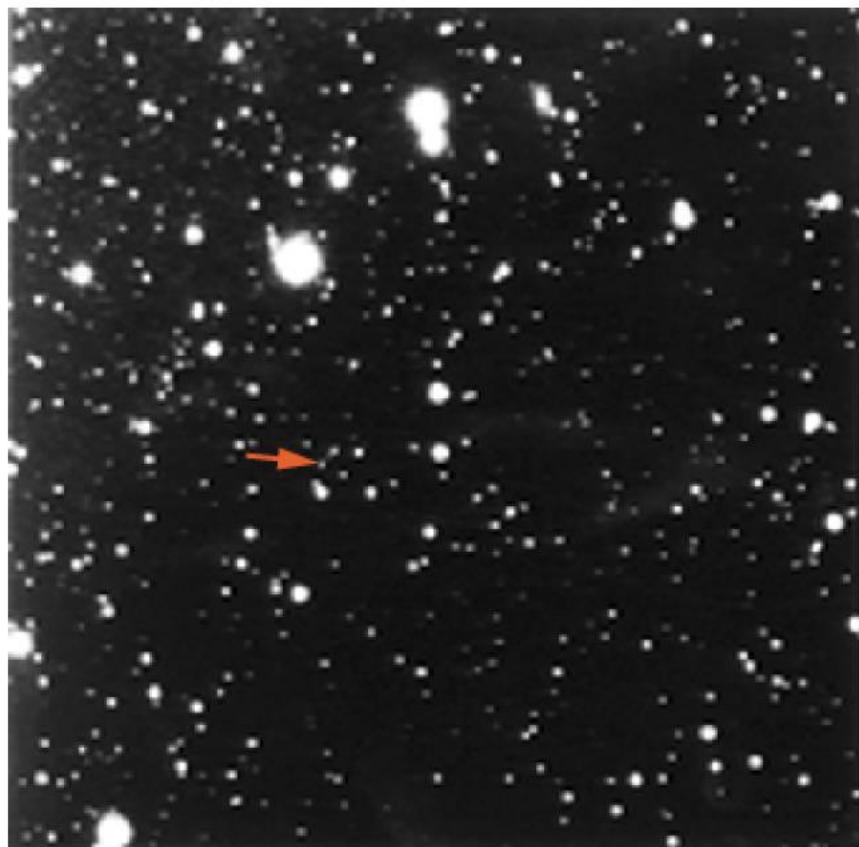


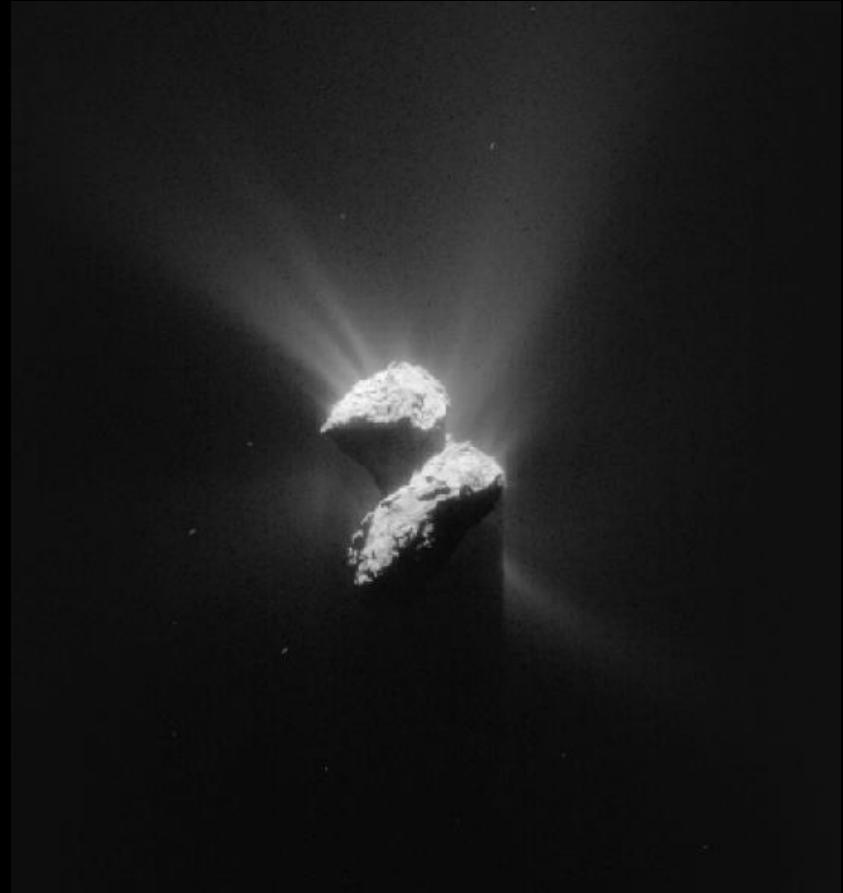
Pluto
2300 km



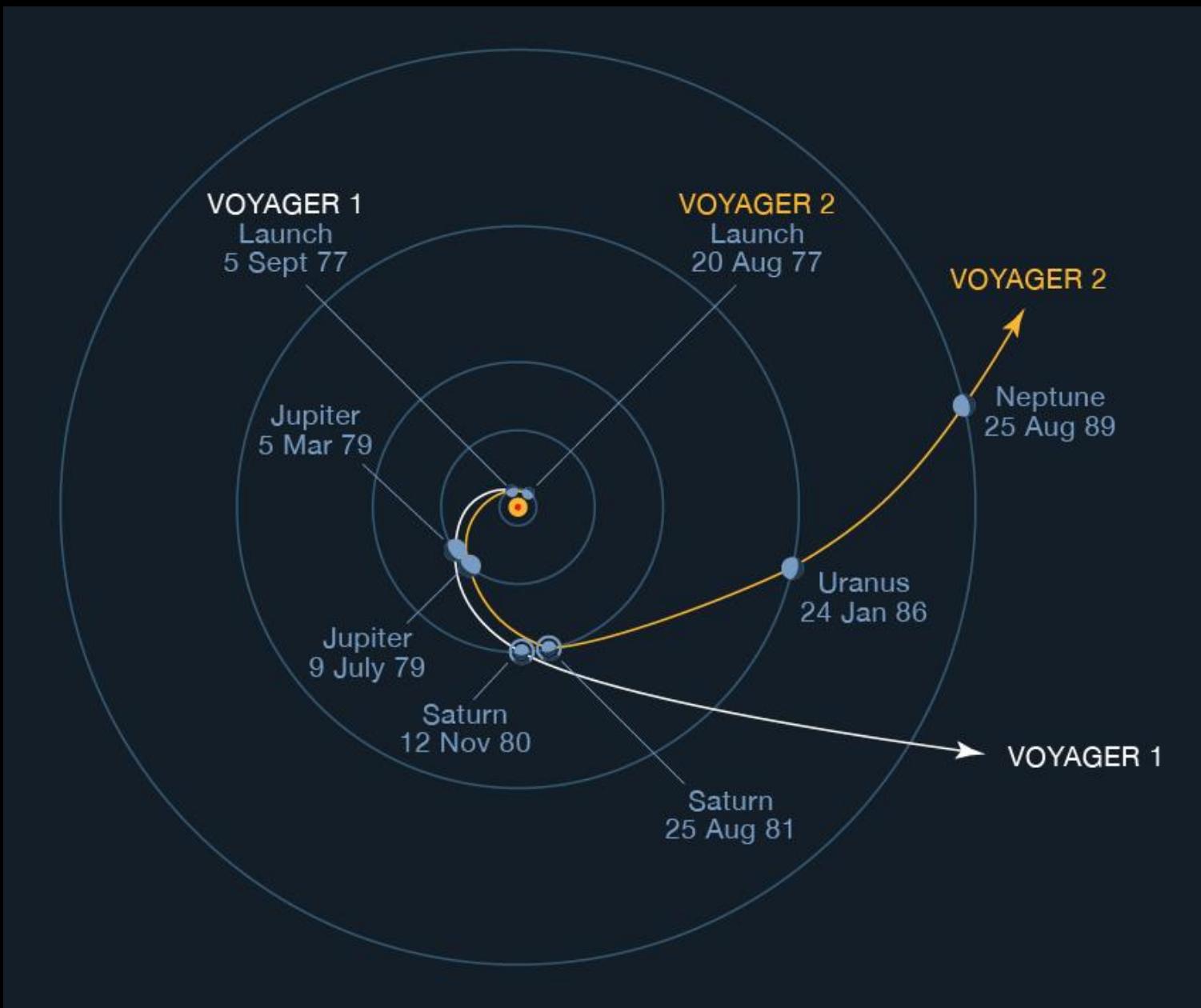
Titania
1580 km

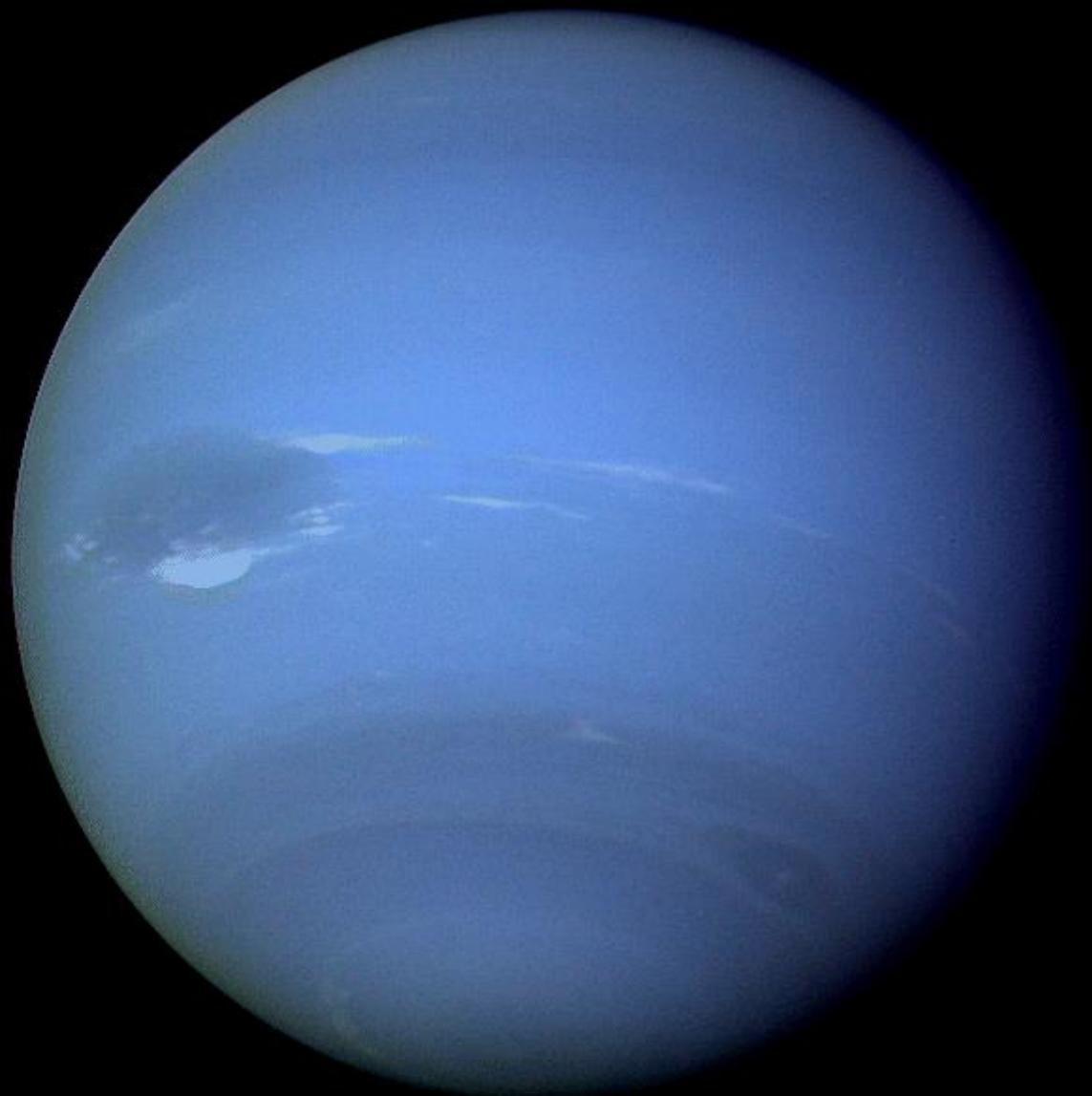
Discovery of Pluto





Asteroids



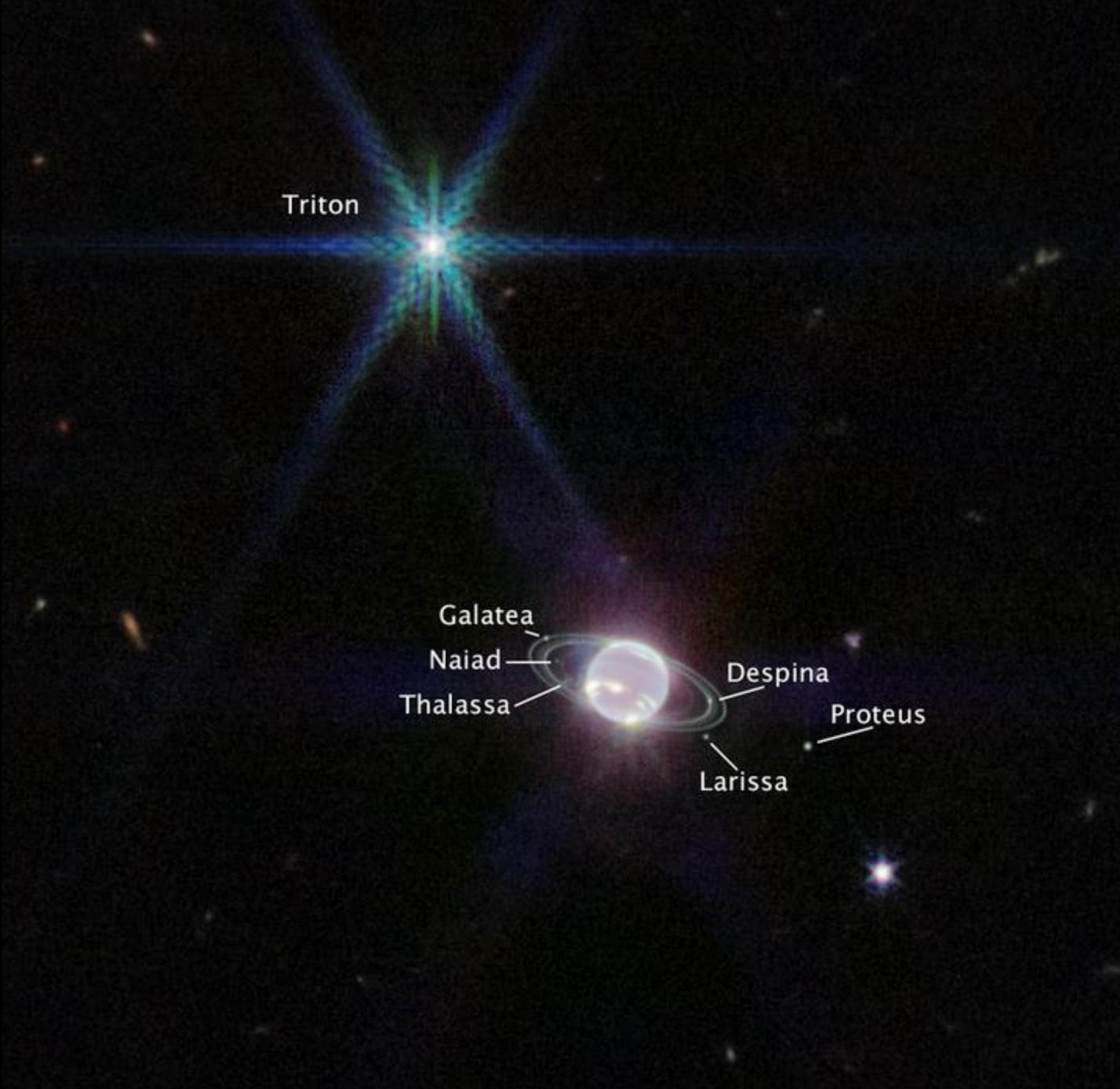


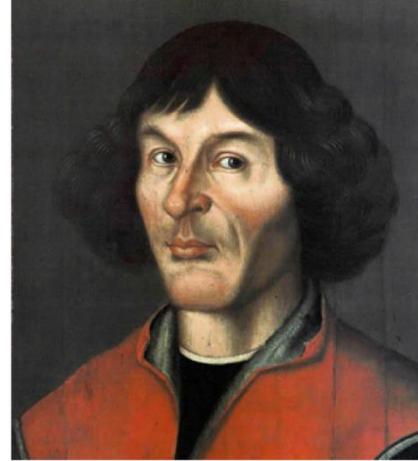
Neptune from Voyager

JAMES WEBB SPACE TELESCOPE

NEPTUNE







Kepler's Laws

(based on Tycho Brahe's data)

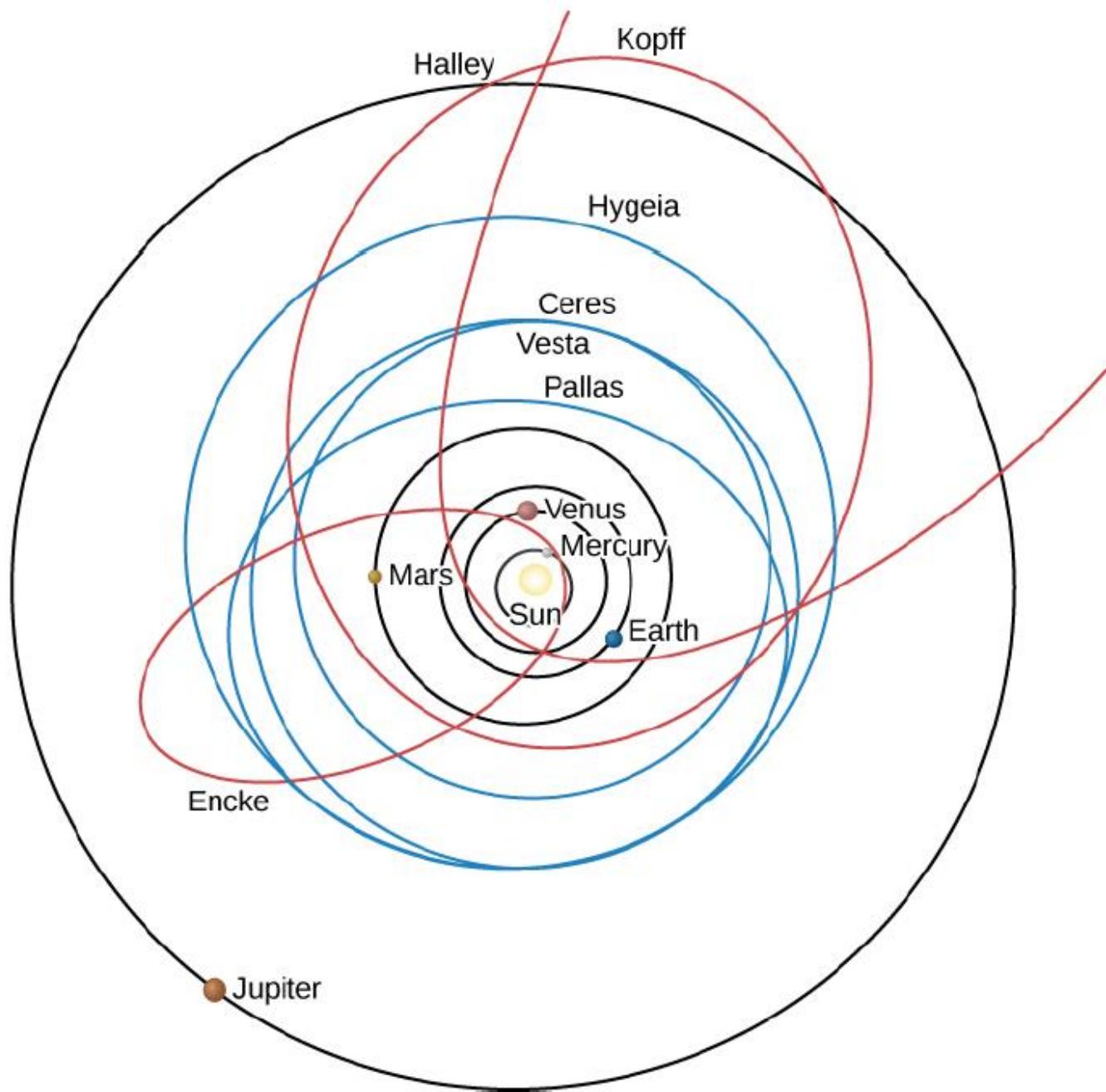


$$P^2 \propto a^3$$

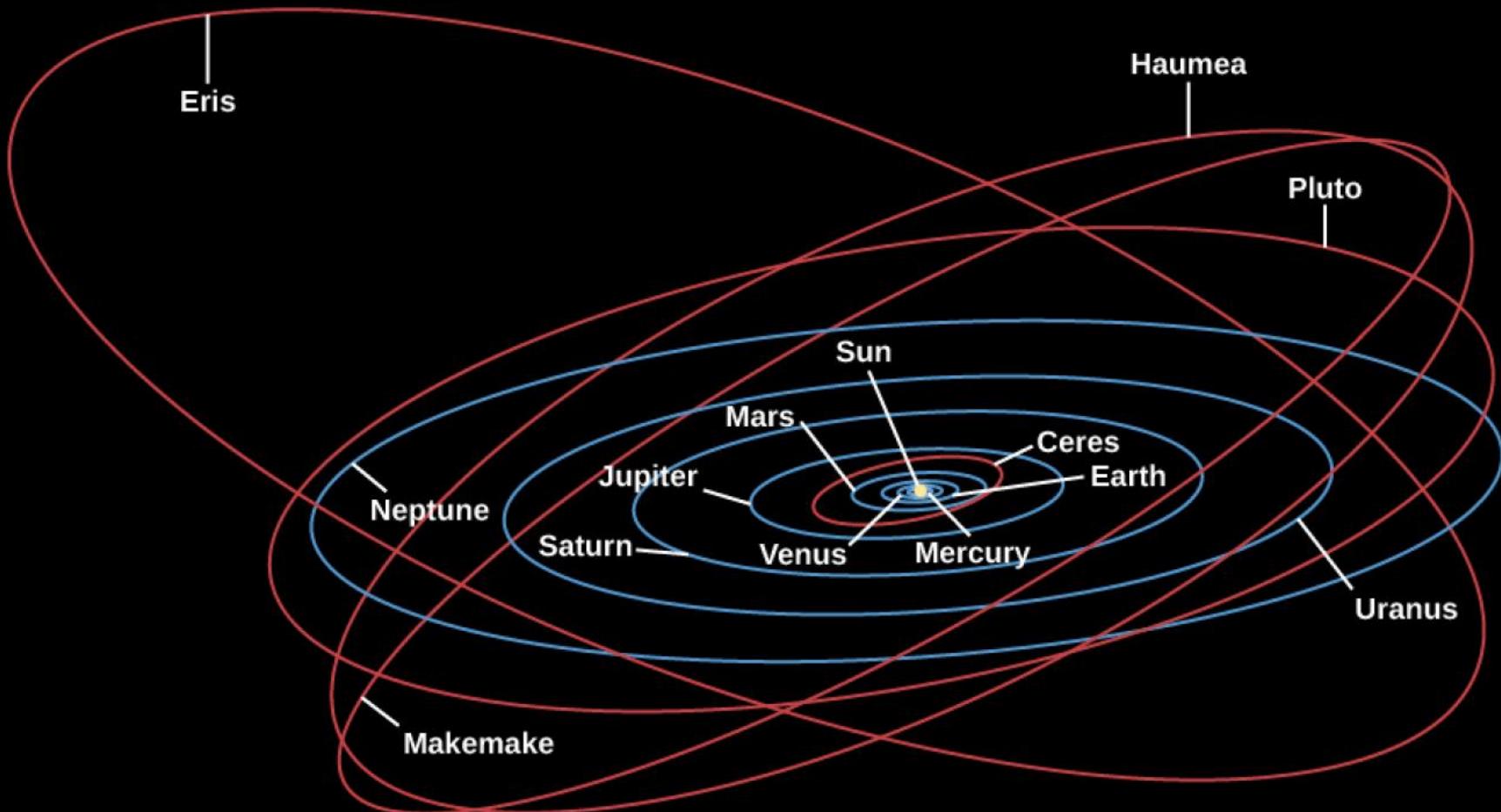
Newton's Laws of Gravity

- **Newton's first law:** Every object will continue to be in a state of rest or move at a constant speed in a straight line unless it is compelled to change by an outside force.
- **Newton's second law:** The change of motion of a body is proportional to and in the direction of the force acting on it.
- **Newton's third law:** For every action there is an equal and opposite reaction (or: the mutual actions of two bodies upon each other are always equal and act in opposite directions).

$$F_{\text{gravity}} = G \frac{M_1 M_2}{R^2}$$



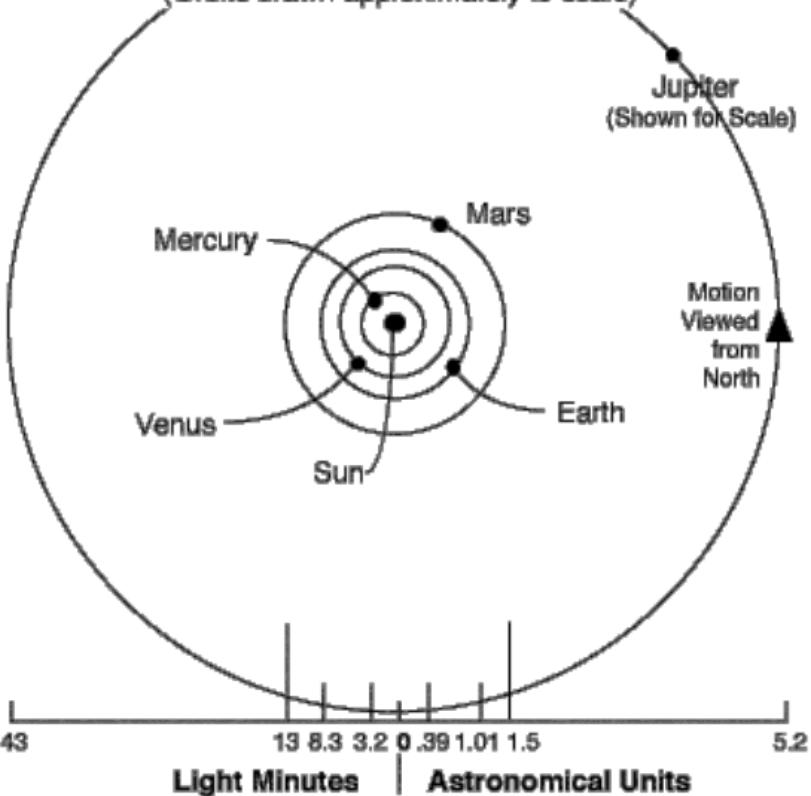
Solar System



The orbits of the planets are very nearly circular and concentric.

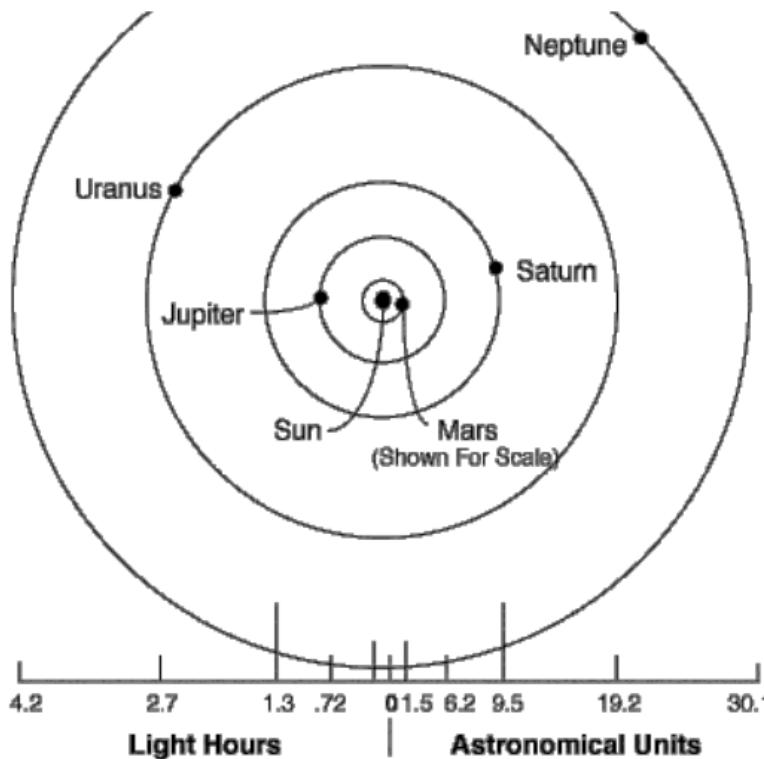
Mean Distances Of The Terrestrial Planets From The Sun

(Orbits drawn approximately to scale)

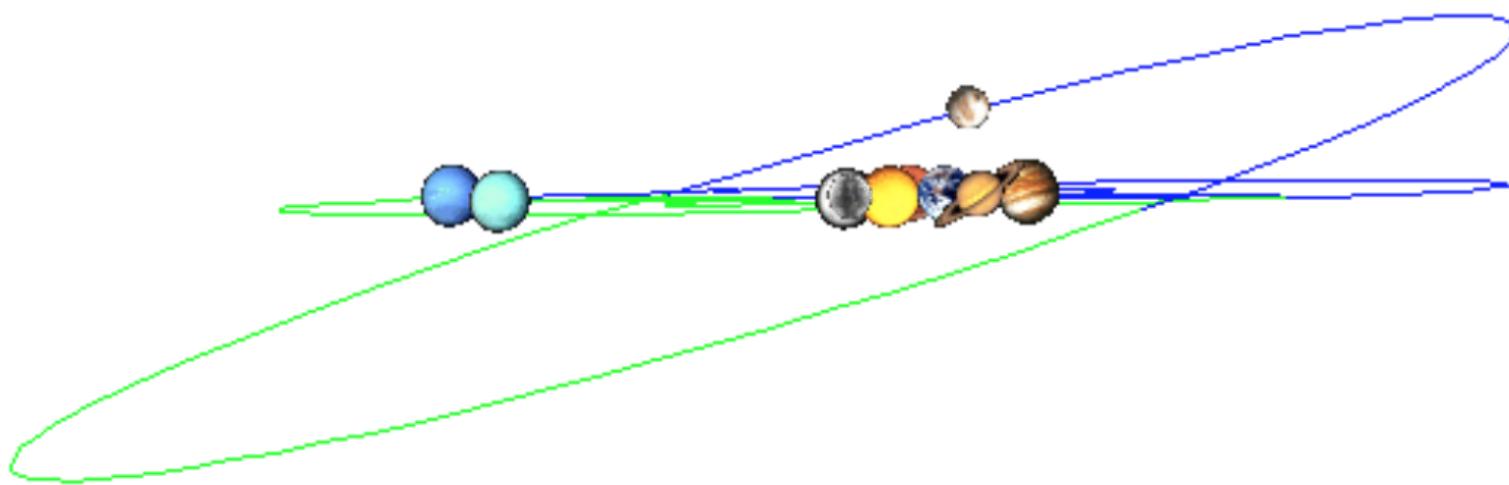


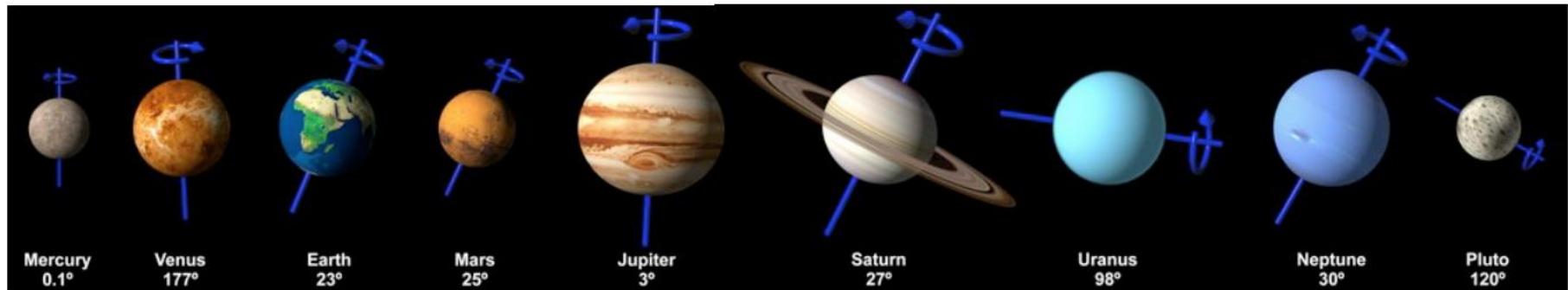
Mean Distances Of The Jovian Planets From The sun

(Orbits drawn approximately to scale.)



All the planets (but not Pluto) orbit in the same direction and in the same plane: the **ecliptic** (to within 6°).





Orbital Data for the Planets

Planet	Semimajor Axis (AU)	Period (y)	Eccentricity
Mercury	0.39	0.24	0.21
Venus	0.72	0.6	0.01
Earth	1	1.00	0.02
Mars	1.52	1.88	0.09
(Ceres)	2.77	4.6	0.08
Jupiter	5.20	11.86	0.05
Saturn	9.54	29.46	0.06
Uranus	19.19	84.01	0.05

$$P^2 \propto a^3$$

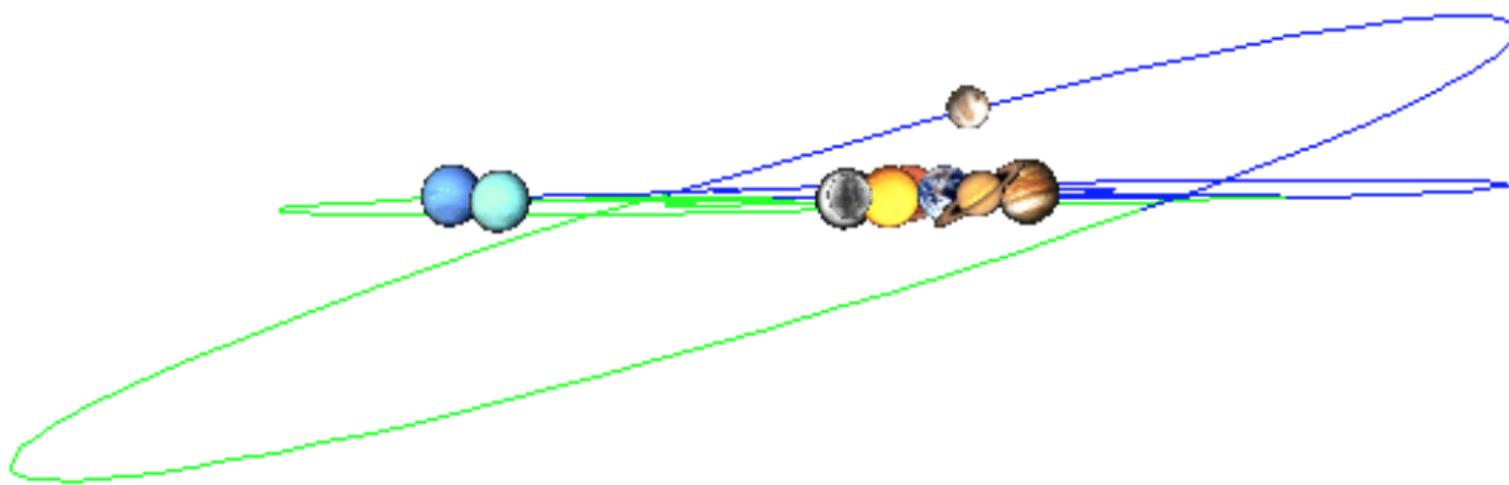


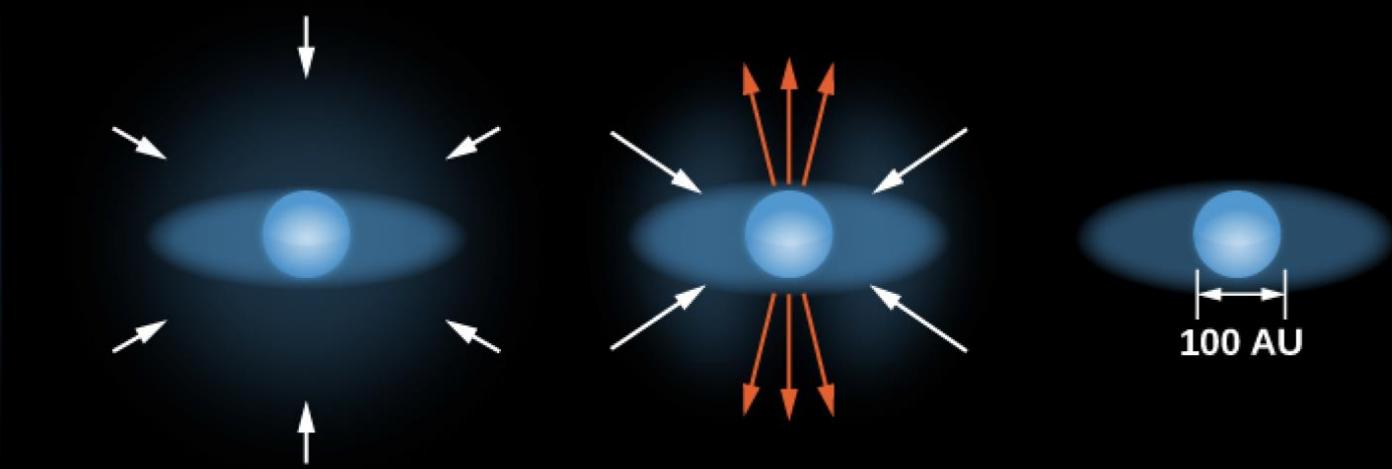
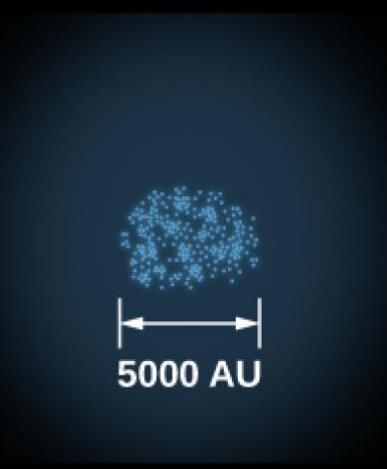
Object	Percentage of Total Mass of Solar System
Sun	99.80
Jupiter	0.10
Comets	0.0005–0.03 (estimate)
All other planets and dwarf planets	0.04
Moons and rings	0.00005

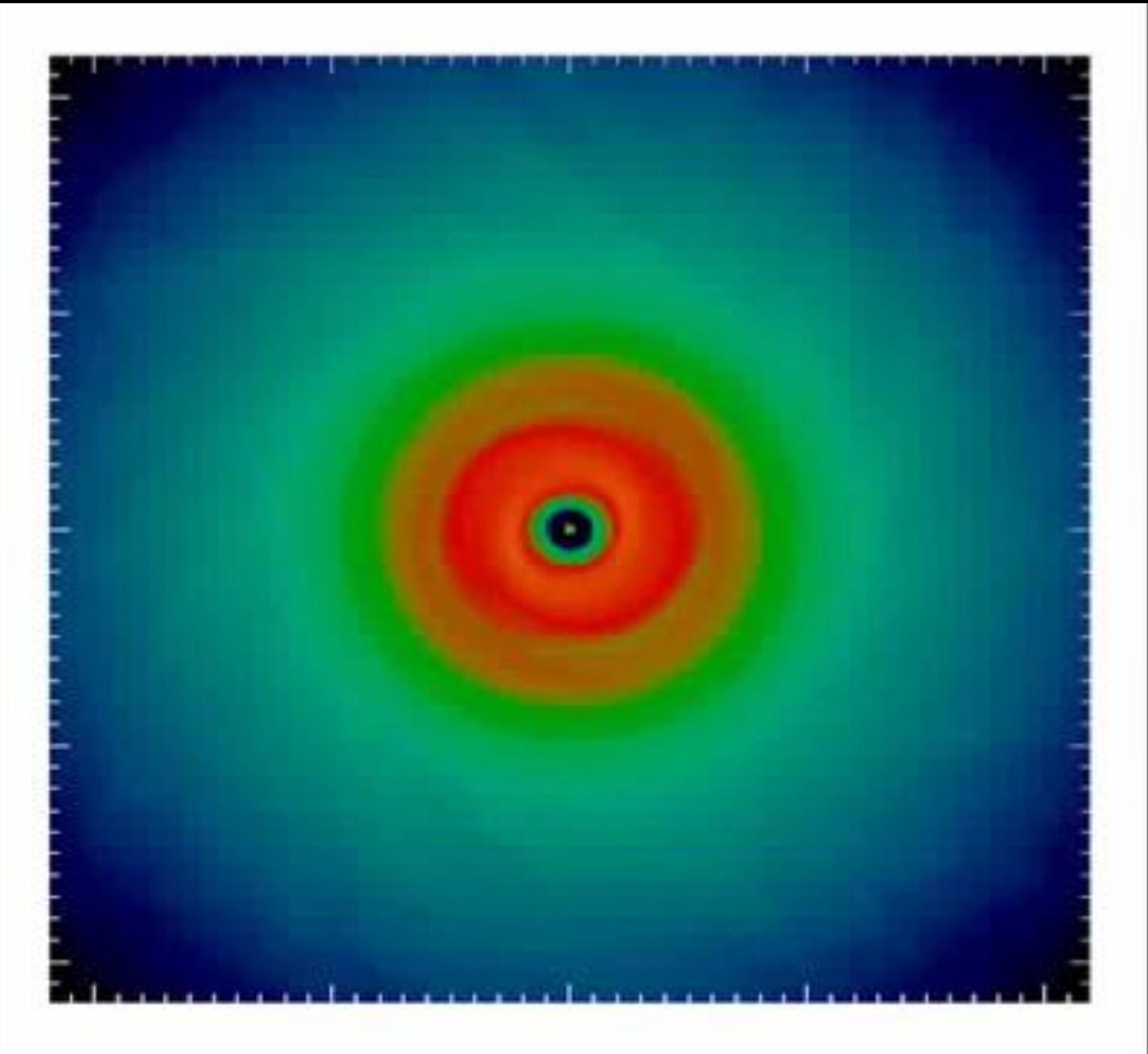
The Planets

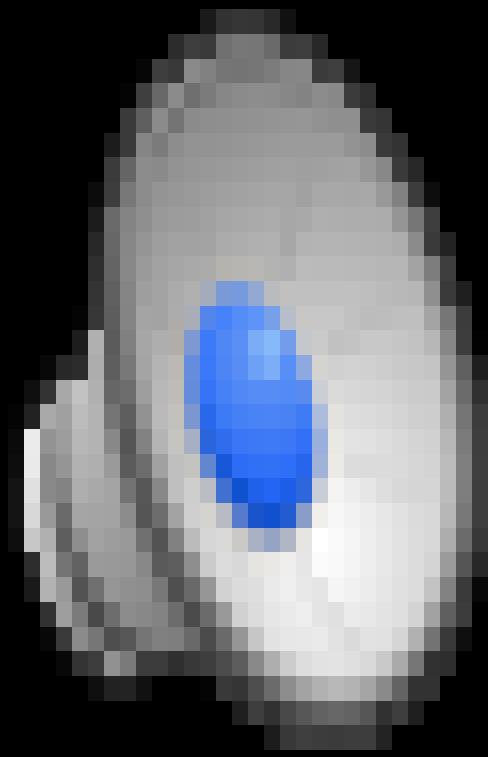
Name	Distance from Sun (AU) ^[2]	Revolution Period (y)	Diameter (km)	Mass (10 ²³ kg)	Density (g/cm ³) ^[3]
Mercury	0.39	0.24	4,878	3.3	5.4
Venus	0.72	0.62	12,120	48.7	5.2
Earth	1.00	1.00	12,756	59.8	5.5
Mars	1.52	1.88	6,787	6.4	3.9
Jupiter	5.20	11.86	142,984	18,991	1.3
Saturn	9.54	29.46	120,536	5686	0.7
Uranus	19.18	84.07	51,118	866	1.3
Neptune	30.06	164.82	49,660	1030	1.6

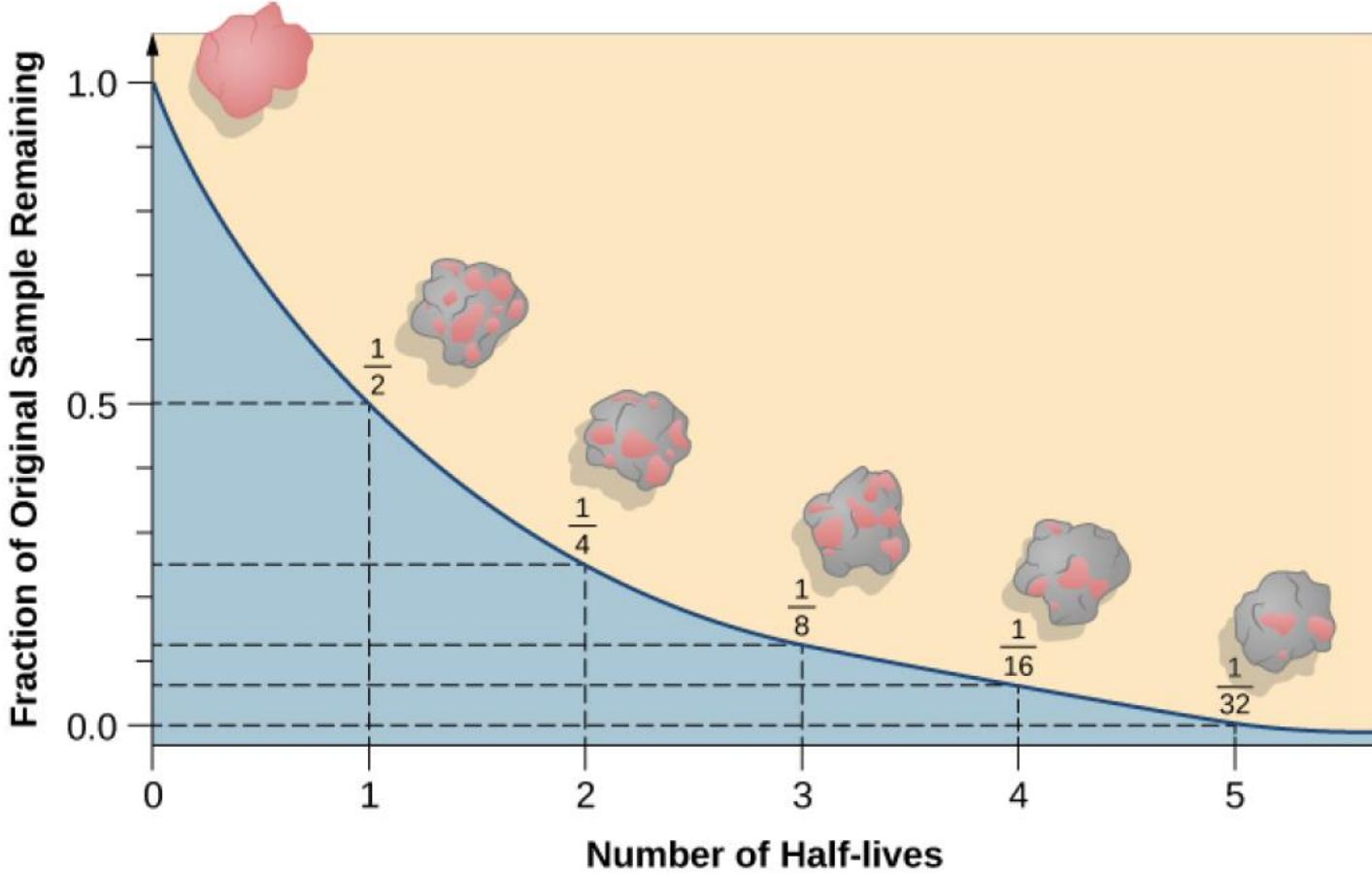
All the planets (but not Pluto) orbit in the same direction and in the same plane: the **ecliptic** (to within 6°).











Parent	Daughter	Half-Life (billions of years)
Samarium-147	Neodymium-143	106
Rubidium-87	Strontium-87	48.8
Thorium-232	Lead-208	14.0
Uranium-238	Lead-206	4.47
Potassium-40	Argon-40	1.31



The terrestrial planets – *rocky worlds*

The terrestrial planets – *rocky worlds*

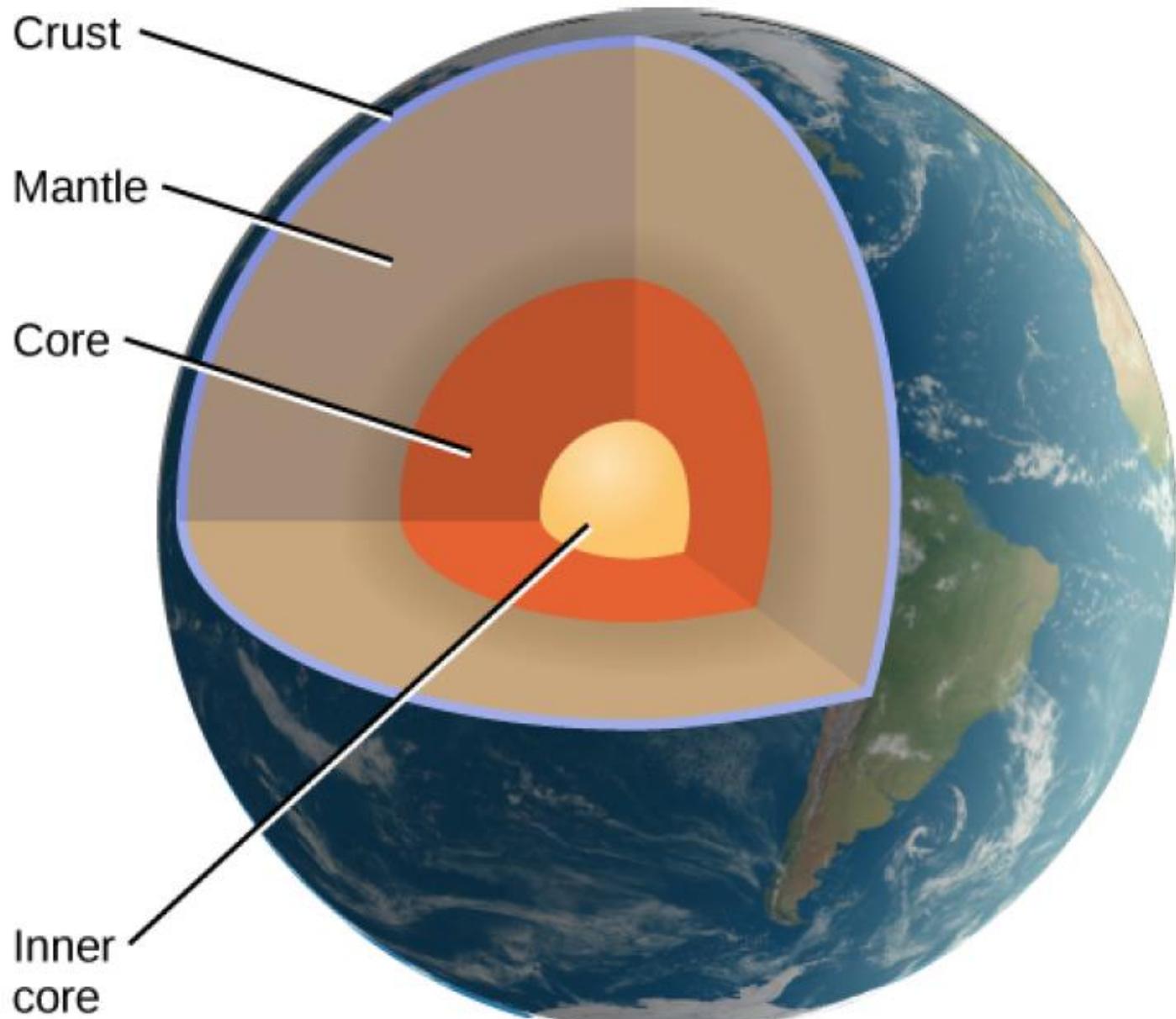


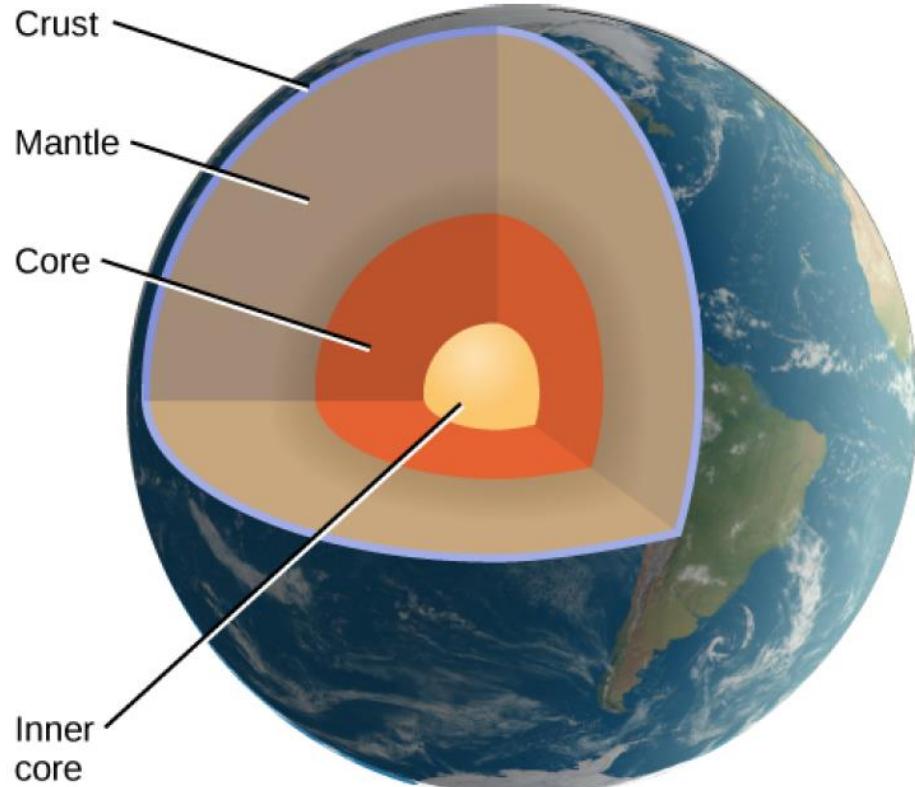
Surfaces of terrestrial planets

Bombardment (collisions with asteroids/comets)

Volcanism (includes earthquakes)
requires molten core

Erosion (if atmosphere)

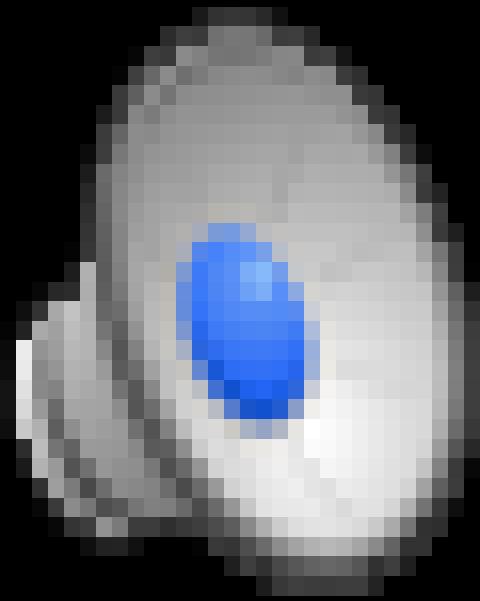


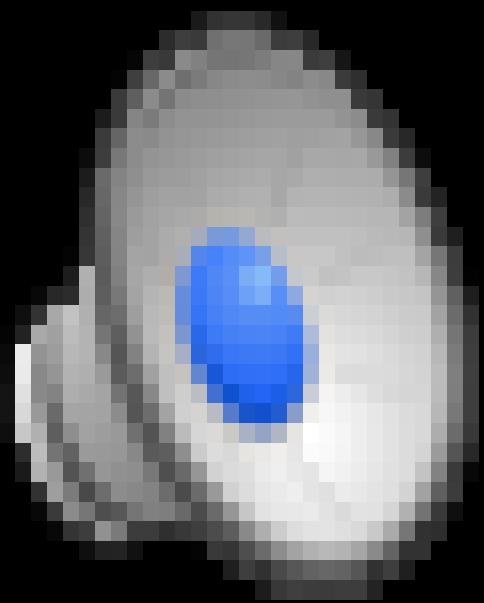


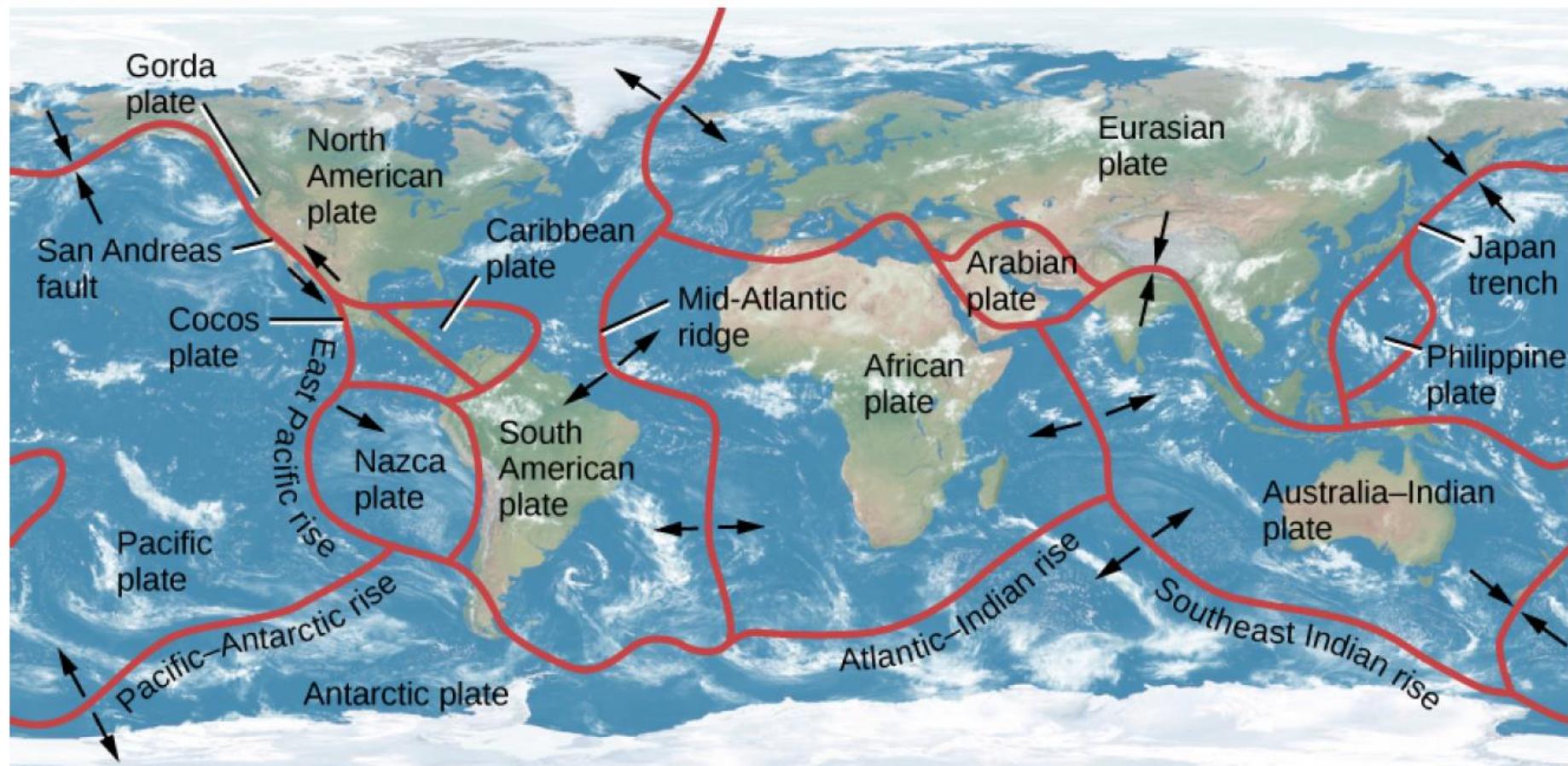
Volcanism:
need molten core

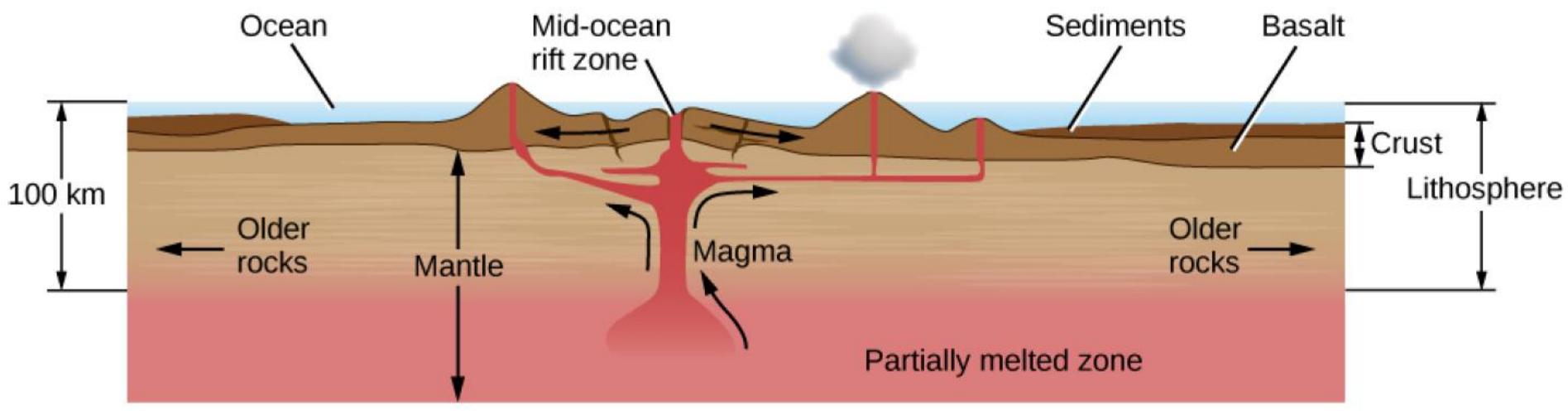
Radioactivity

Larger planets have
cores that stay
molten for longer

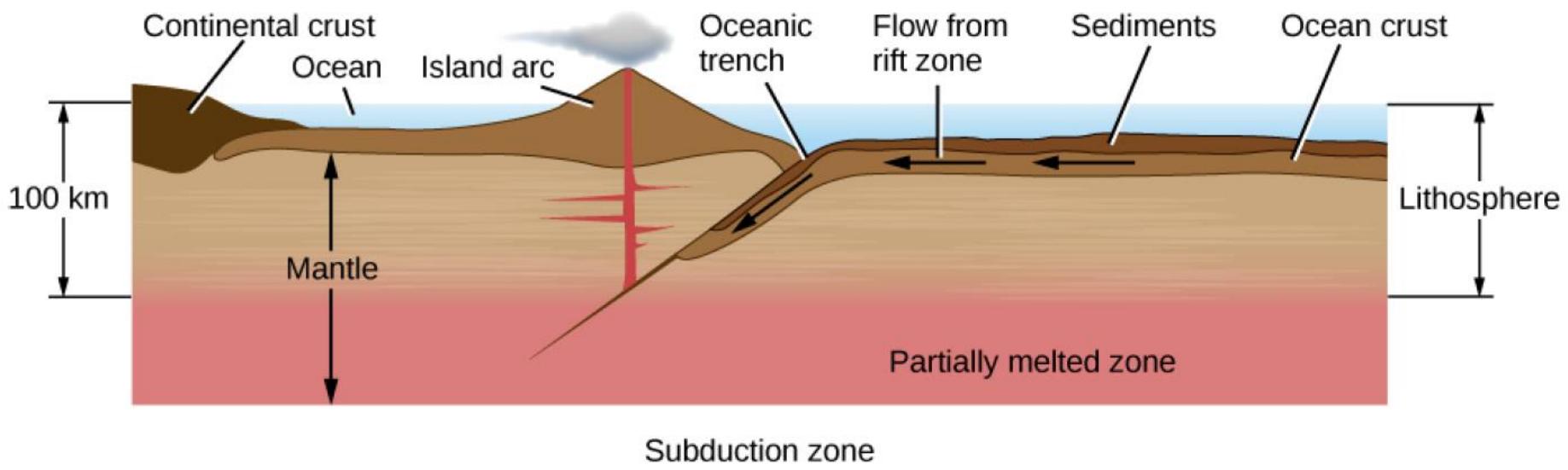








Rift zone

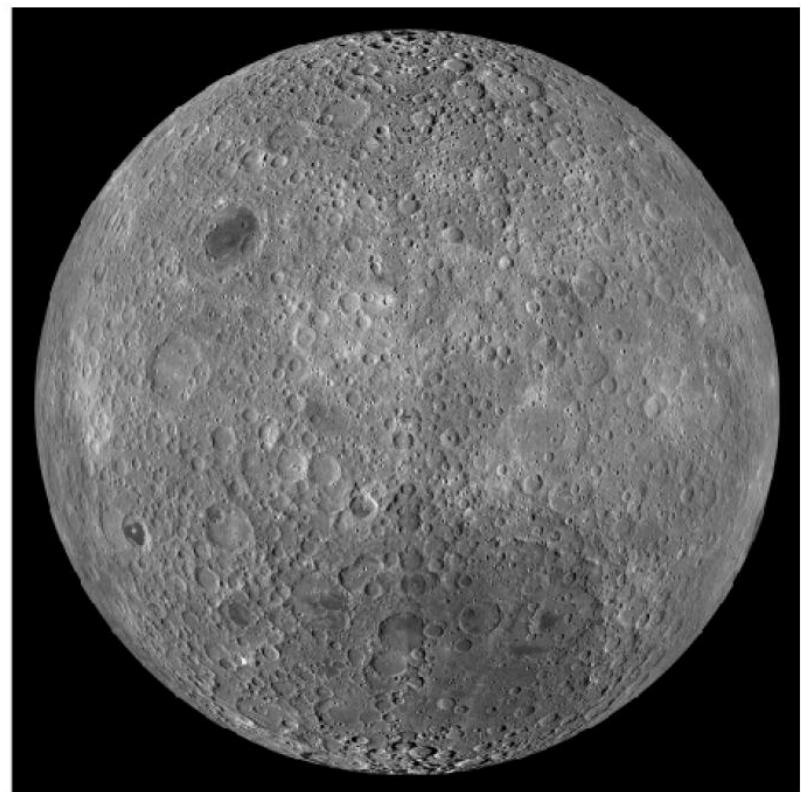
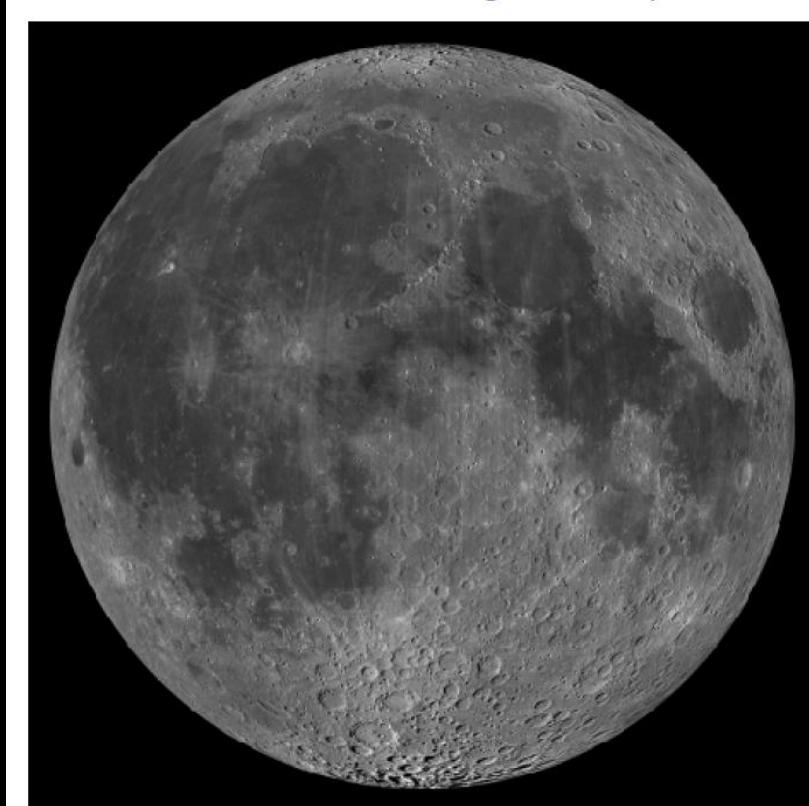
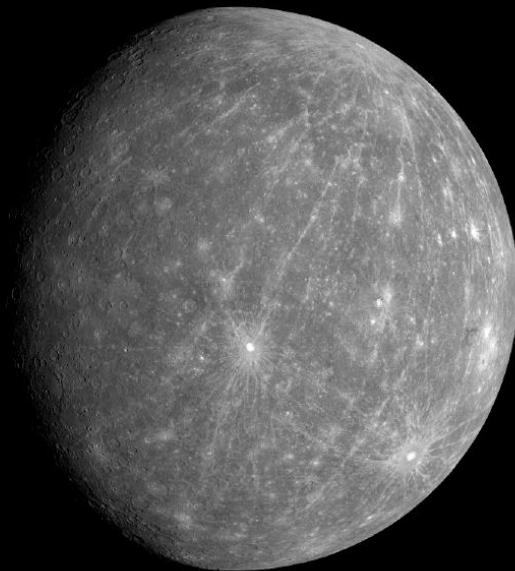


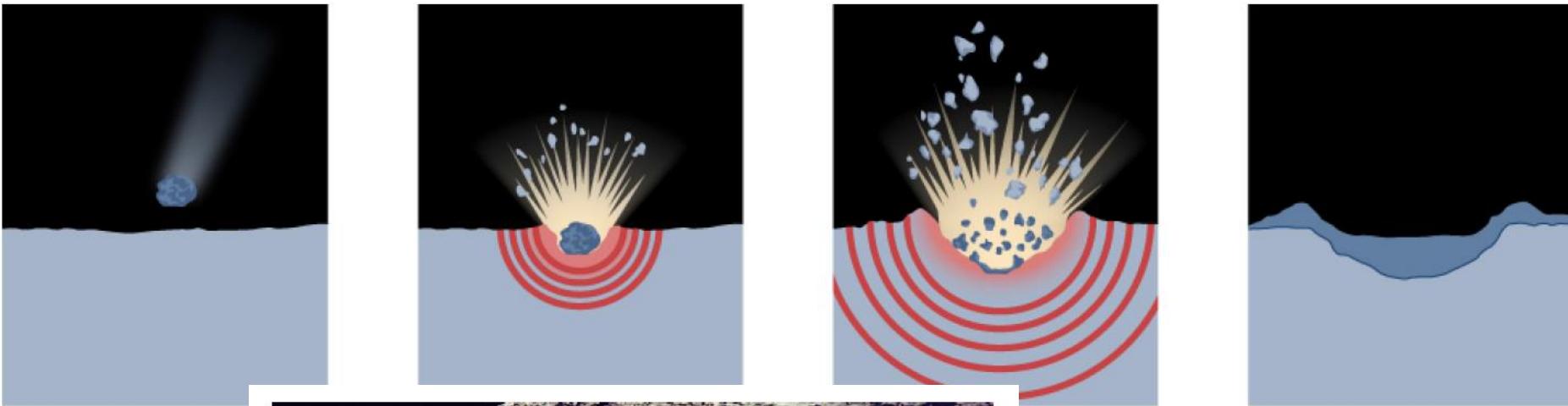
Subduction zone

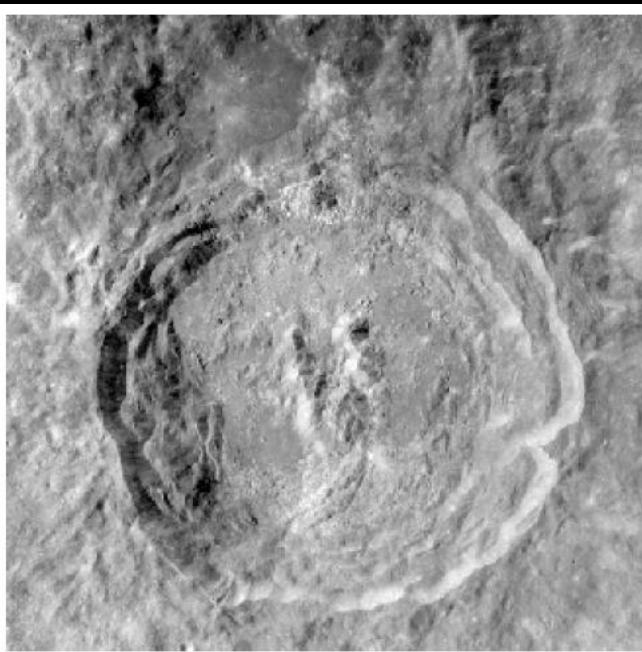
Properties of the Moon and Mercury

Property	Moon	Mercury
Mass (Earth = 1)	0.0123	0.055
Diameter (km)	3476	4878
Density (g/cm ³)	3.3	5.4
Surface gravity (Earth = 1)	0.17	0.38
Escape velocity (km/s)	2.4	4.3
Rotation period (days)	27.3	58.65

The moon and Mercury



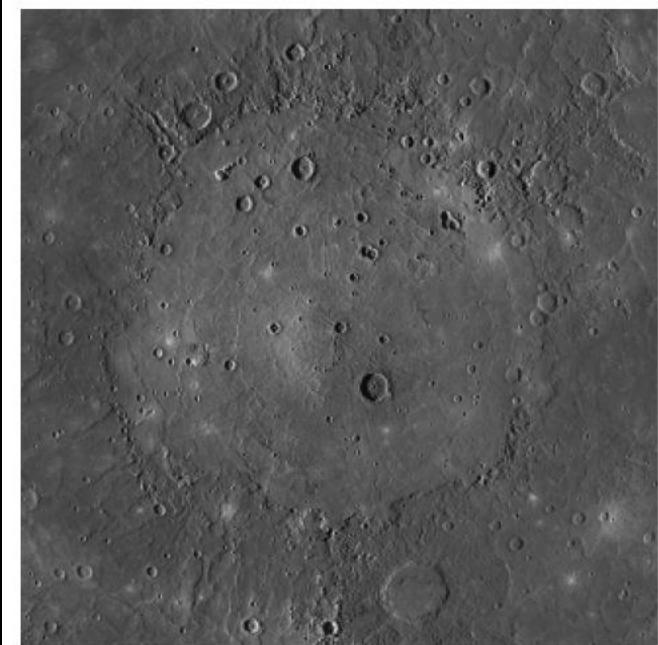


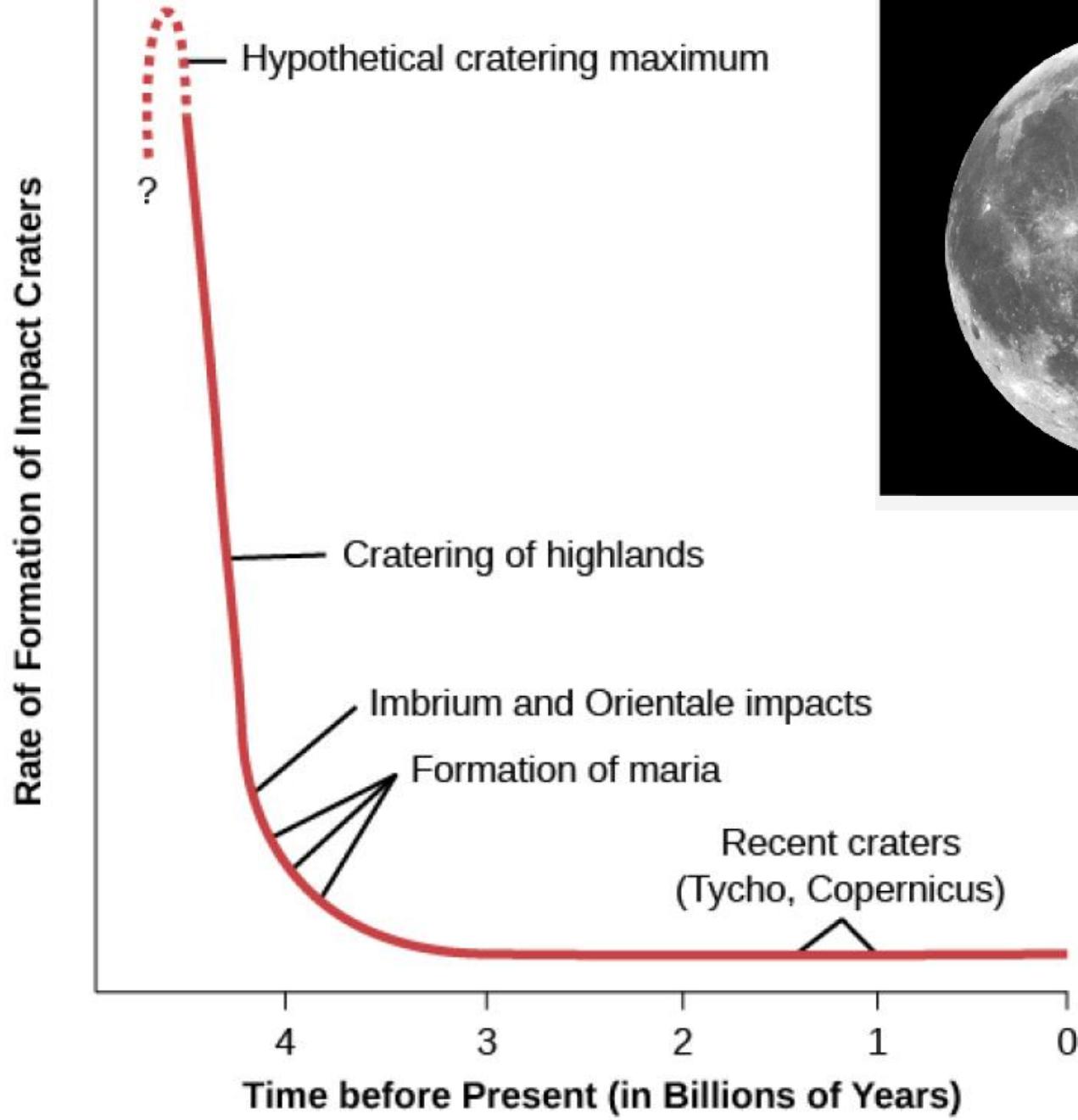


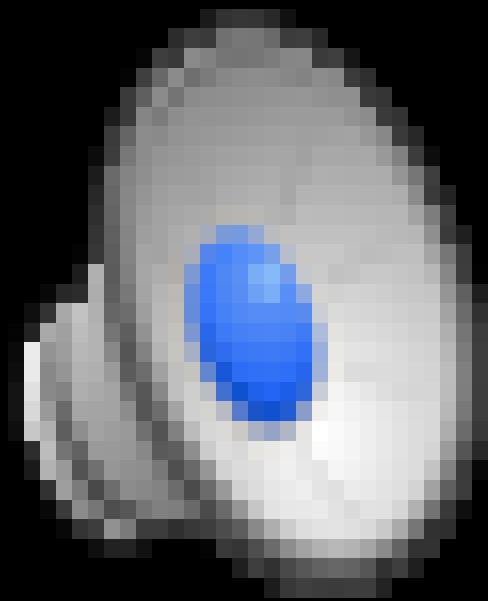
Craters: old surface
Smooth: new surface

Mare=sea









Properties of Earth, Venus, and Mars

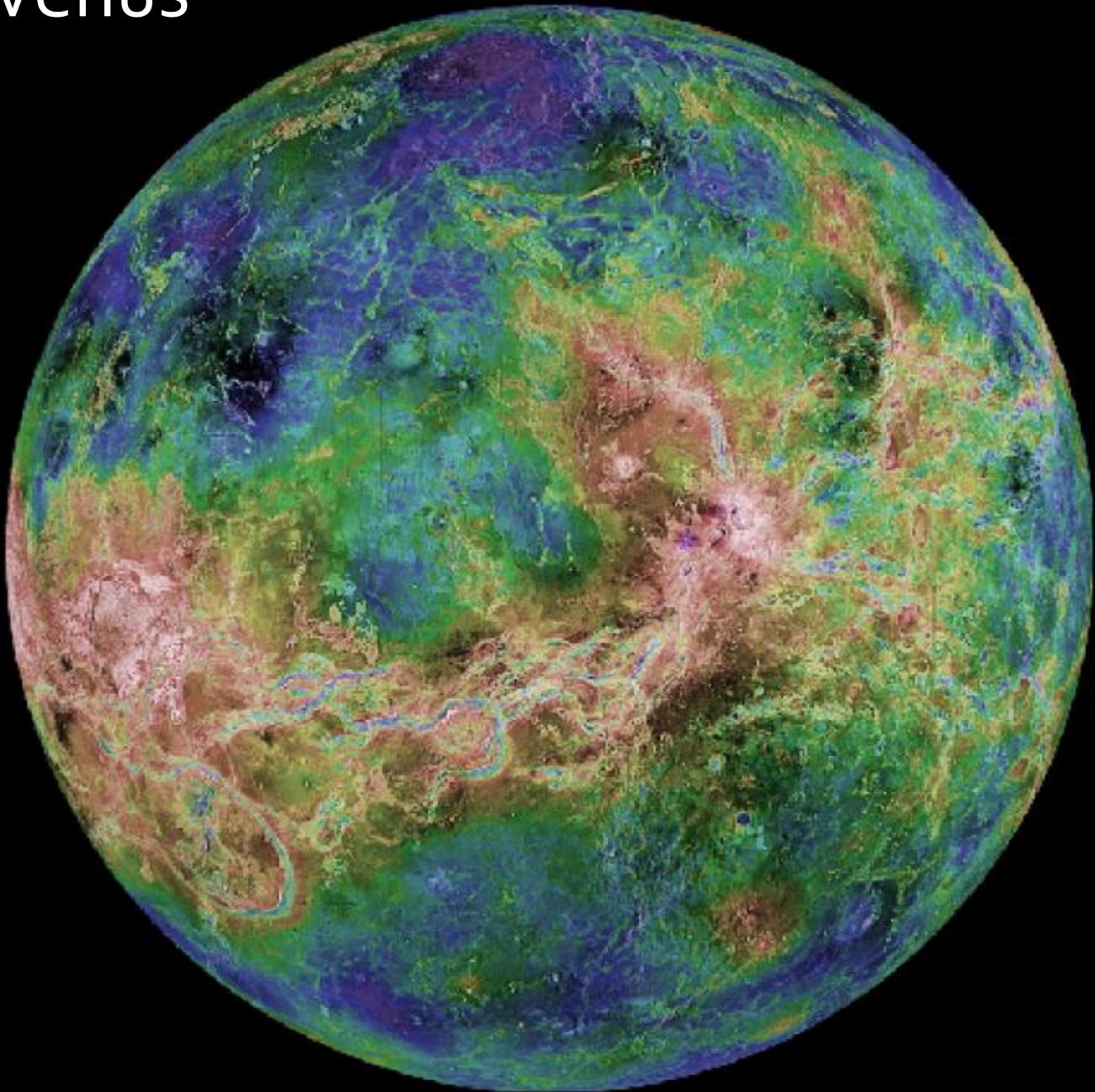
Property	Earth	Venus	Mars
Semimajor axis (AU)	1.00	0.72	1.52
Period (year)	1.00	0.61	1.88
Mass (Earth = 1)	1.00	0.82	0.11
Diameter (km)	12,756	12,102	6,790
Density (g/cm ³)	5.5	5.3	3.9
Surface gravity (Earth = 1)	1.00	0.91	0.38
Escape velocity (km/s)	11.2	10.4	5.0
Rotation period (hours or days)	23.9 h	243 d	24.6 h
Surface area (Earth = 1)	1.00	0.90	0.28
Atmospheric pressure (bar)	1.00	90	0.007

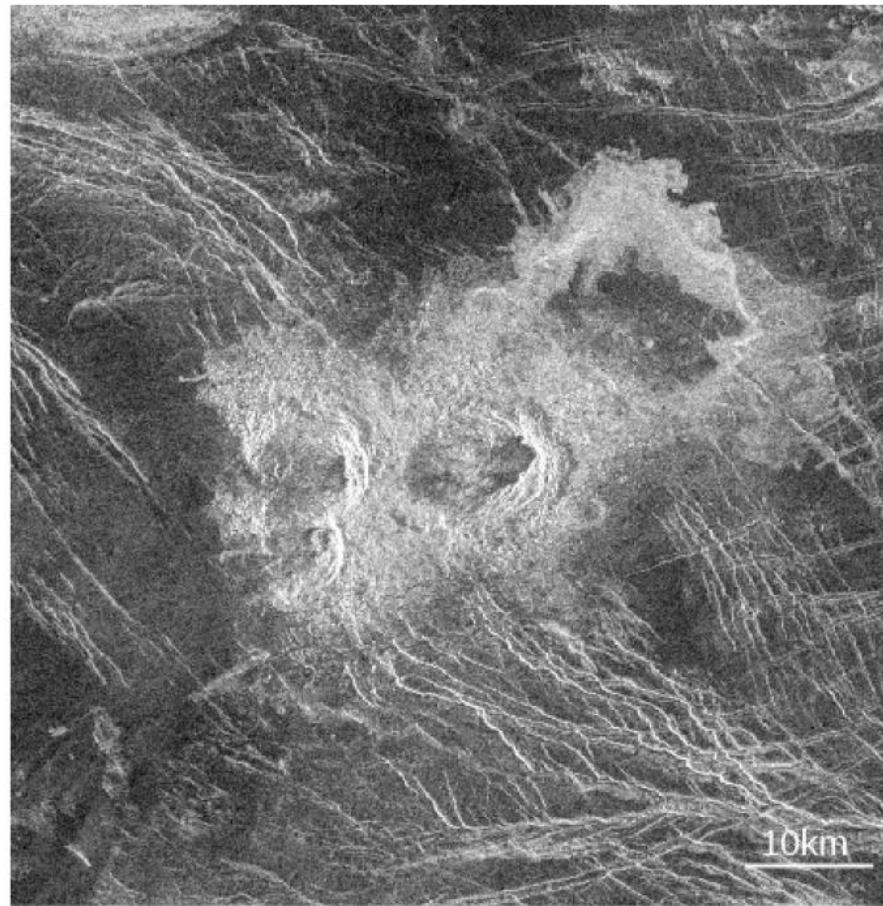
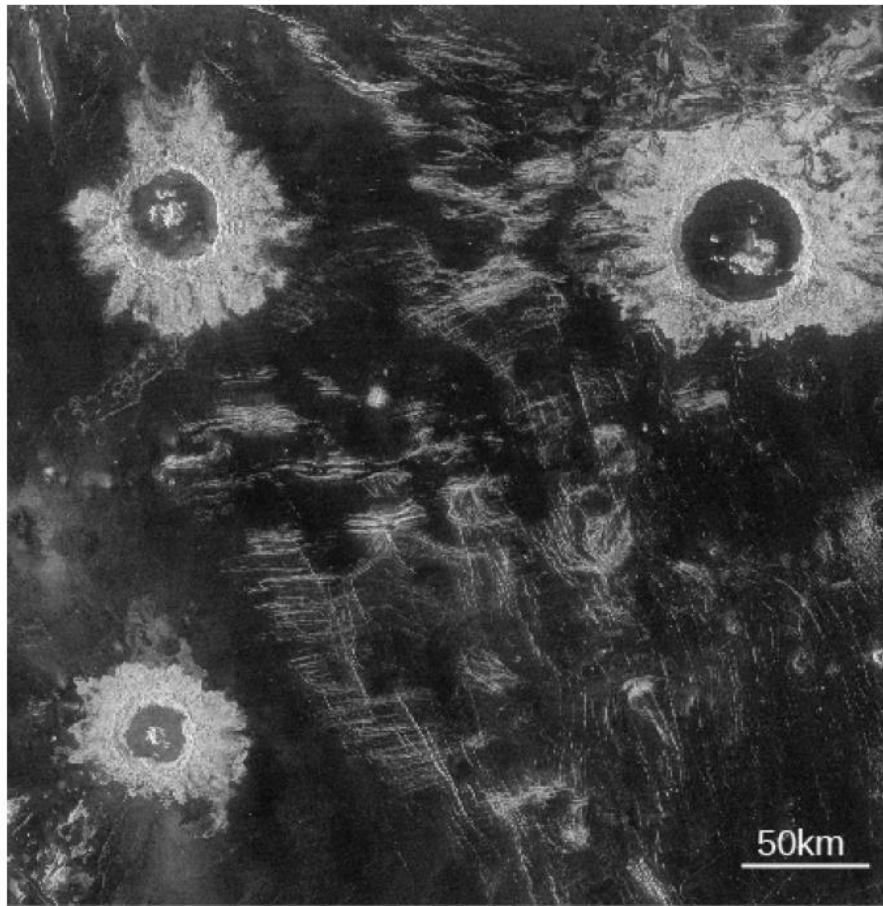
Venus

Thick atmosphere

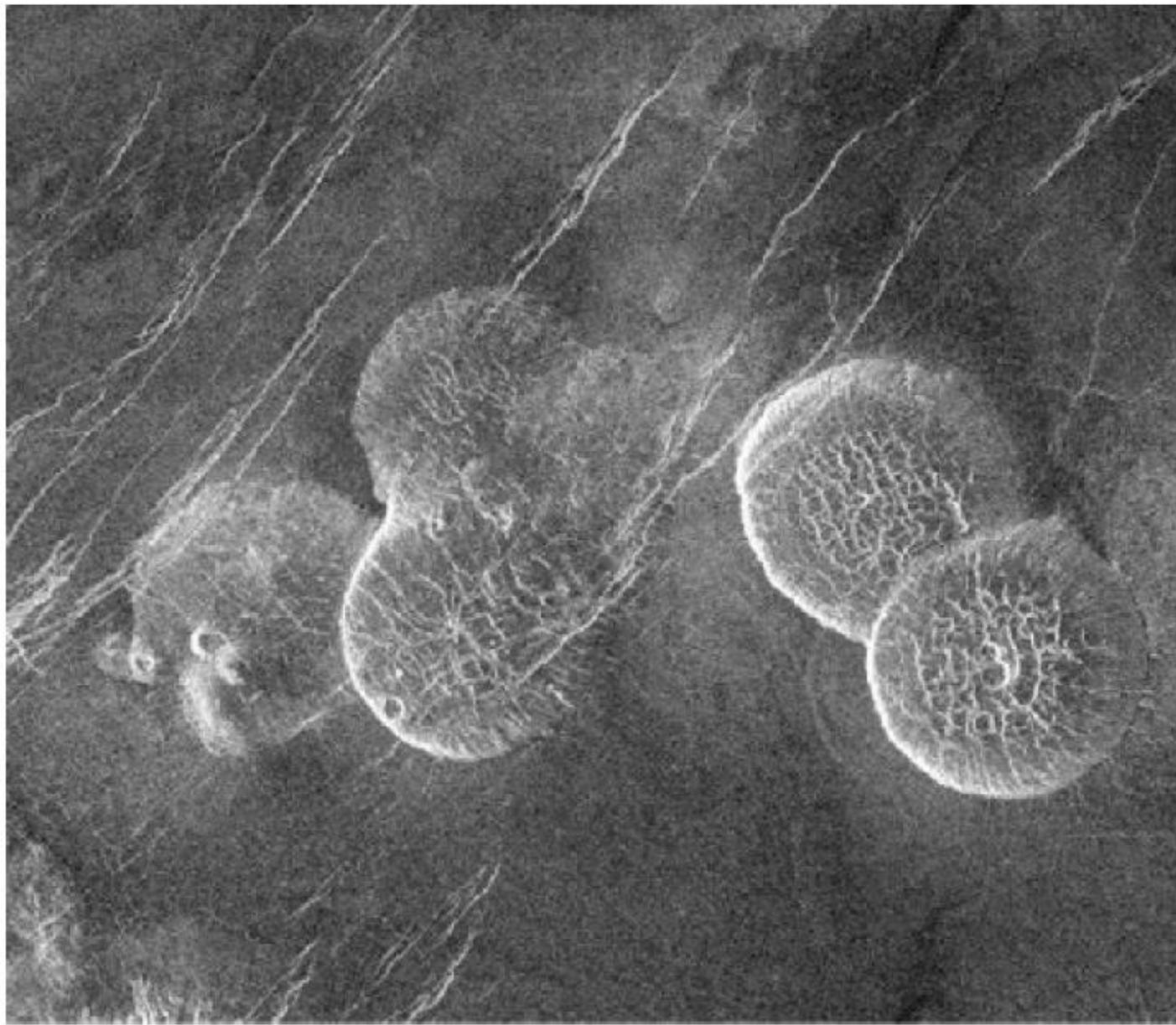


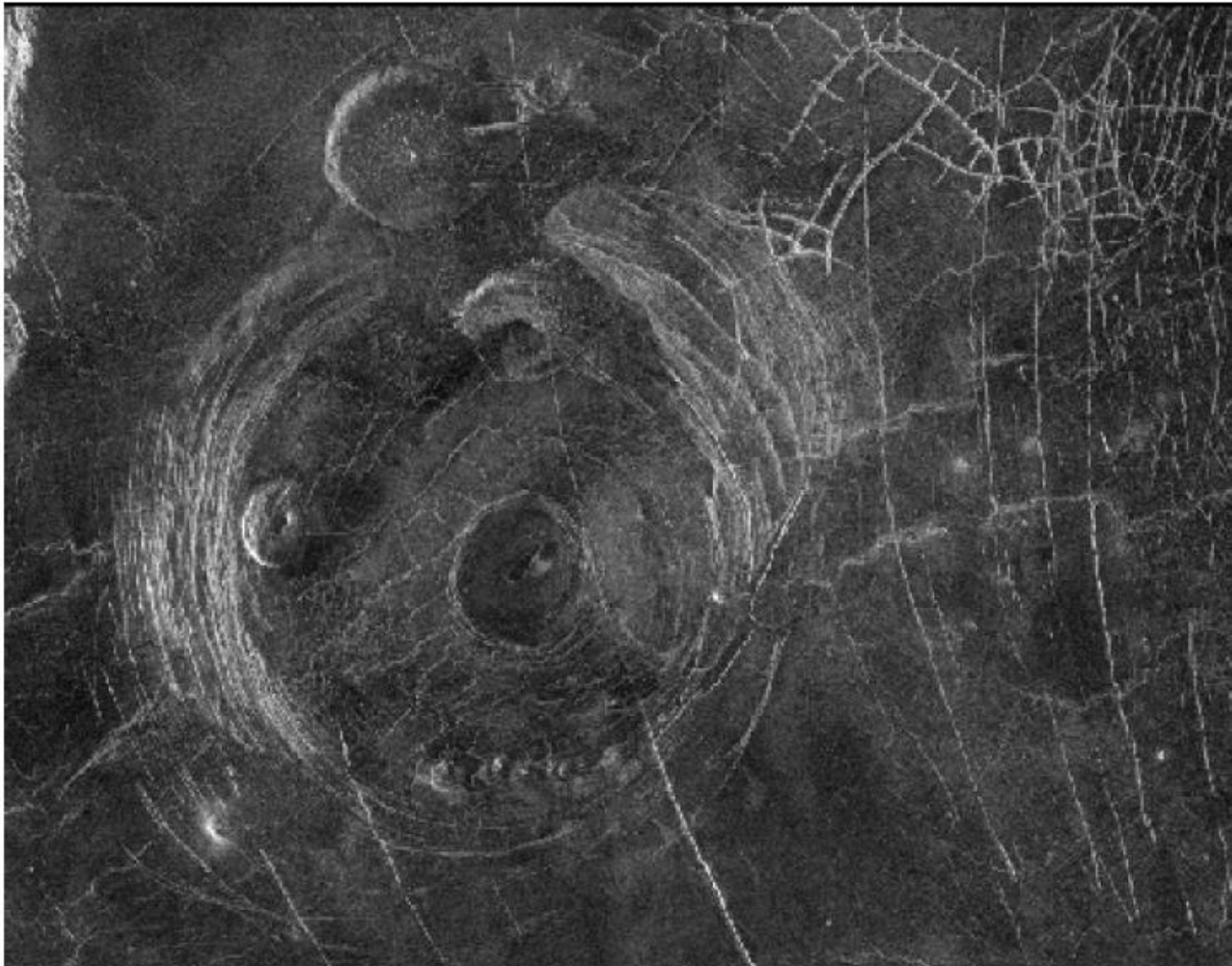
Venus

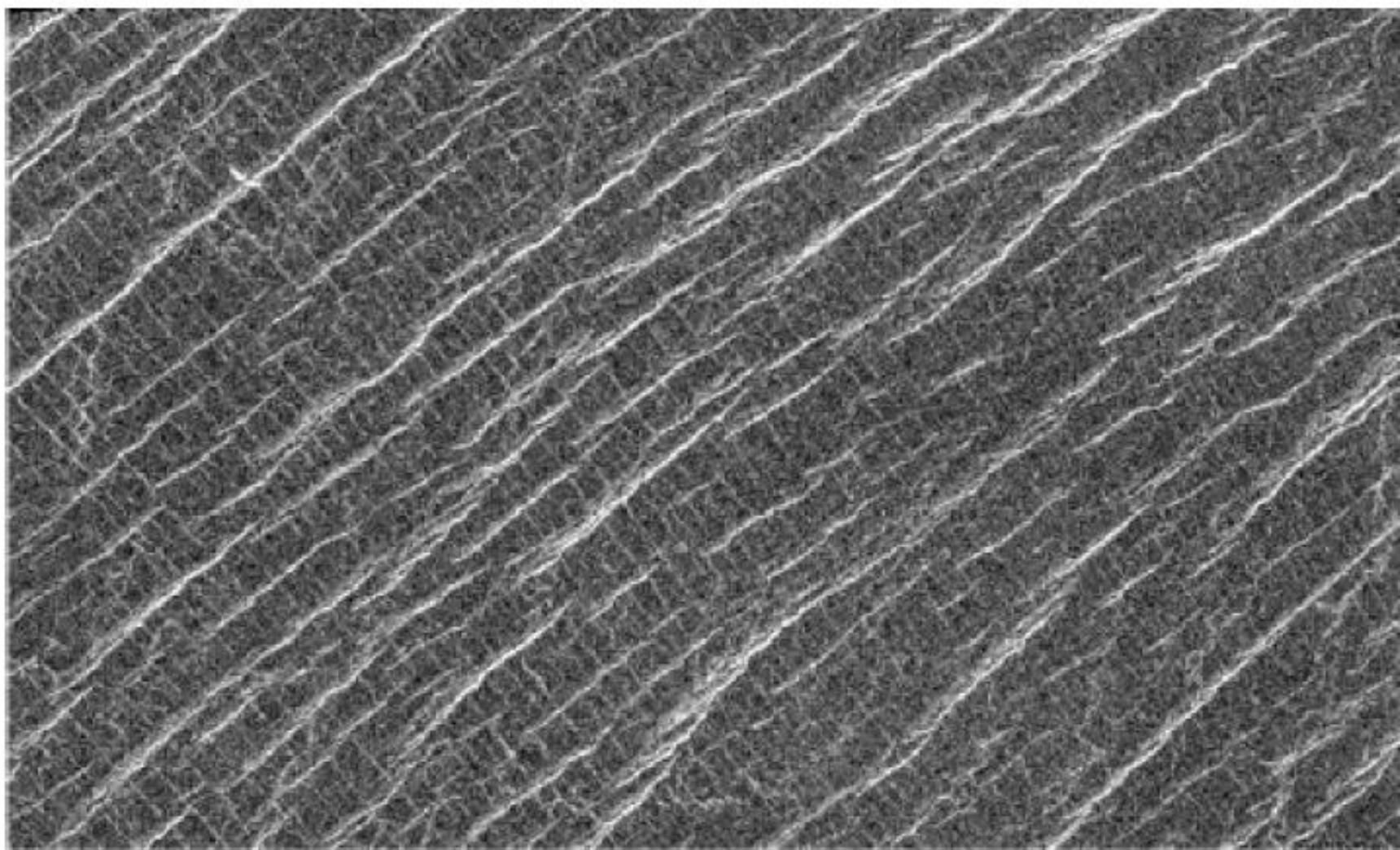


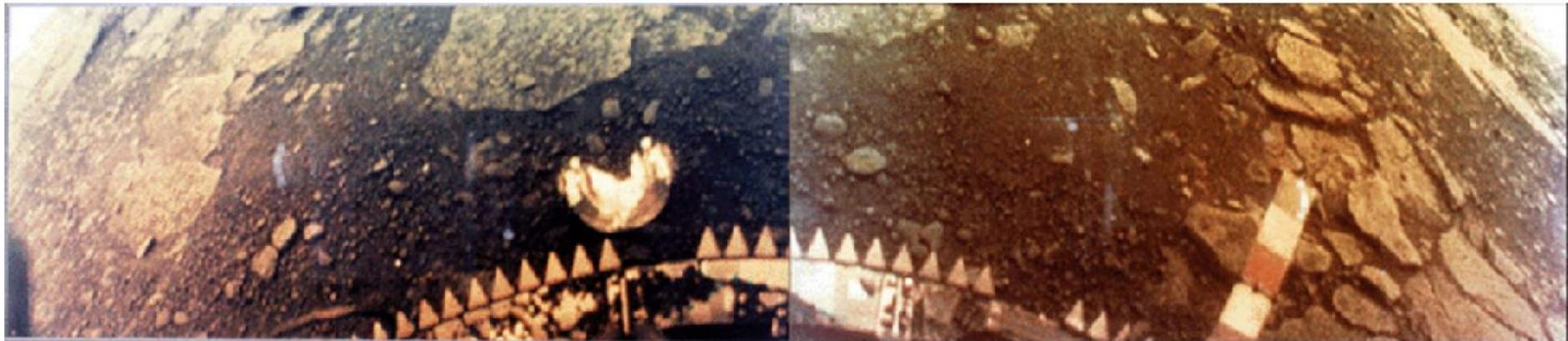


Only a few landers
(all Soviet Union)



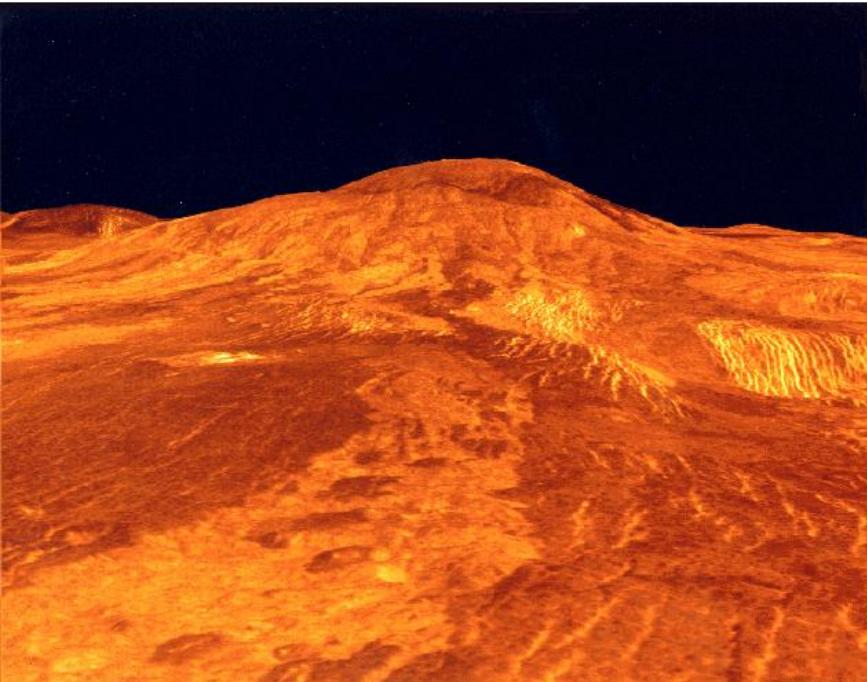






Venus has hardly any impact craters; the surface density of craters indicates most of the surface is only 600 million years old; but craters do not appear to be eroding. Where are all the older craters?

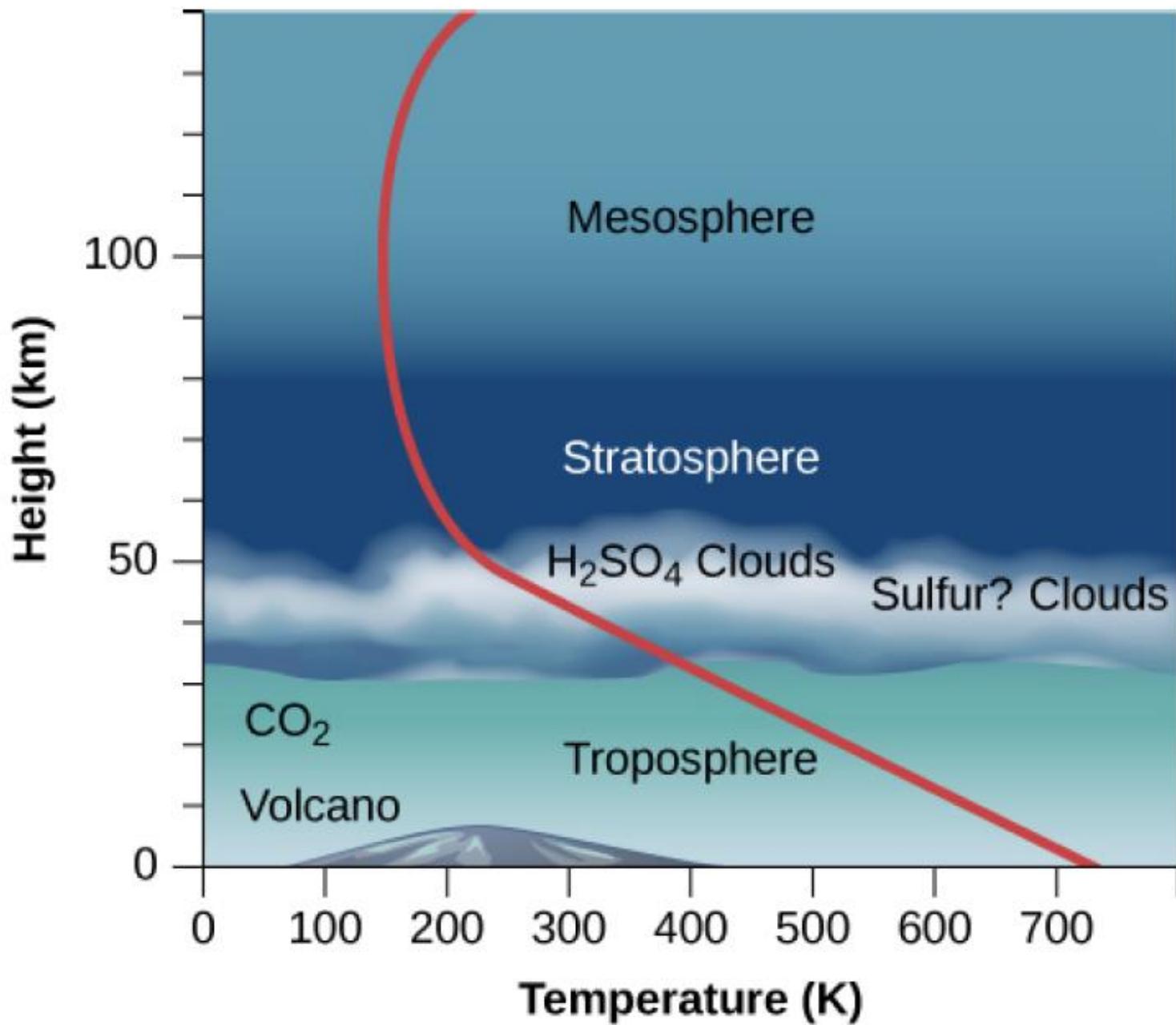
Perhaps Venus undergoes periodic **catastrophic resurfacing**. The last such event would have taken place about 600 million years ago.



The volcano Sif Mons. is about 2 km high and nearly 300 km across. There appear to be recent lava flows at the front of the image: these flows are about 120 km long, which suggests that these lavas were also very fluid.

Atmospheric Composition of Earth, Venus, and Mars

Gas	Earth	Venus	Mars
Carbon dioxide (CO ₂)	0.03%	96%	95.3%
Nitrogen (N ₂)	78.1%	3.5%	2.7%
Argon (Ar)	0.93%	0.006%	1.6%
Oxygen (O ₂)	21.0%	0.003%	0.15%
Neon (Ne)	0.002%	0.001%	0.0003%

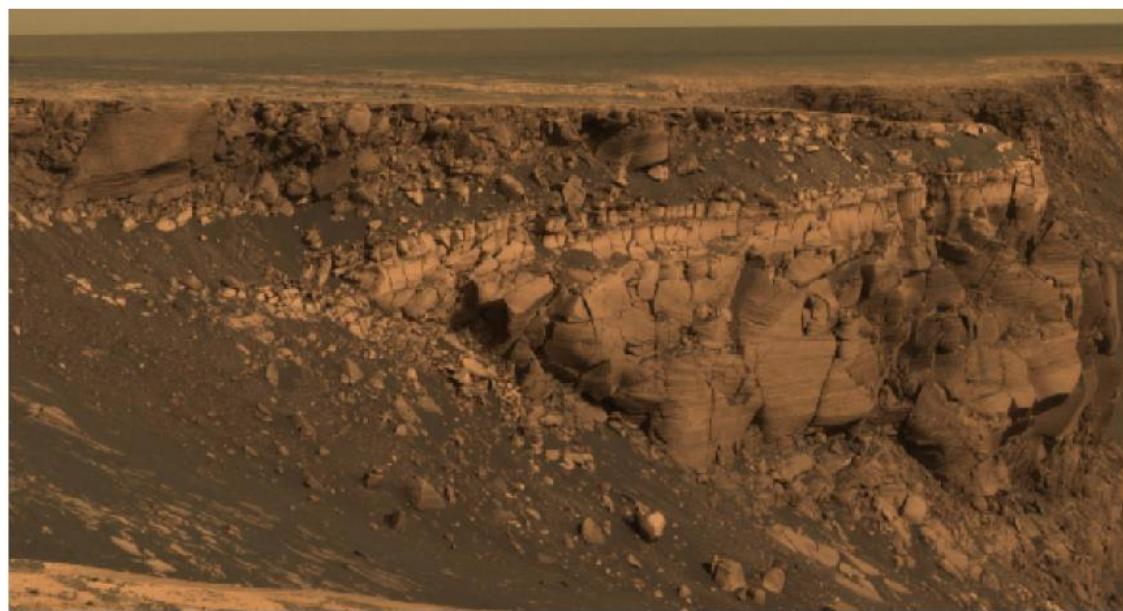


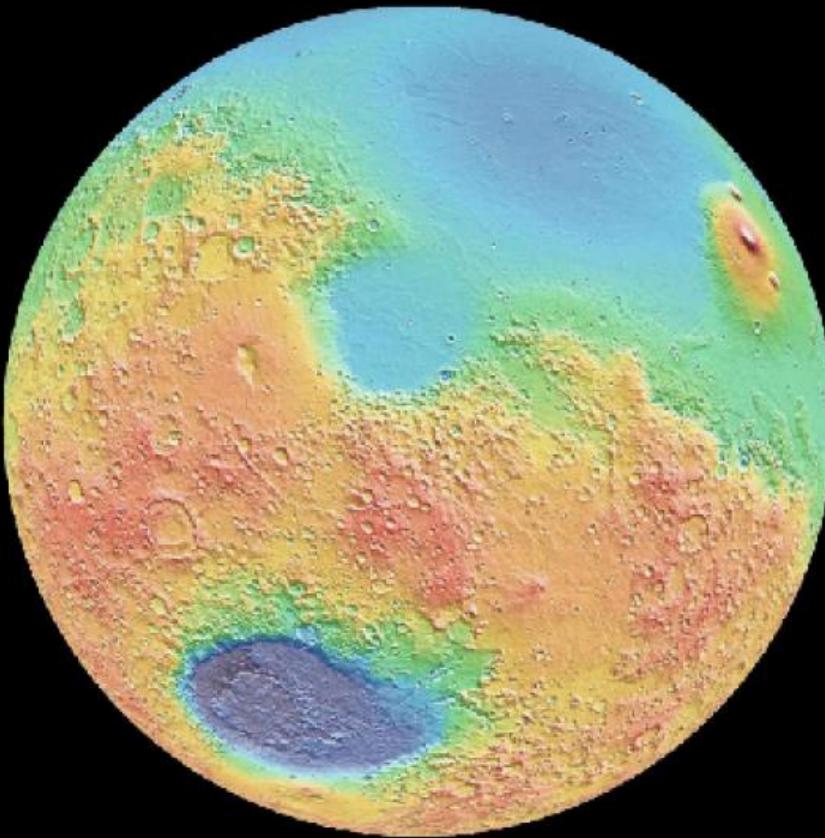
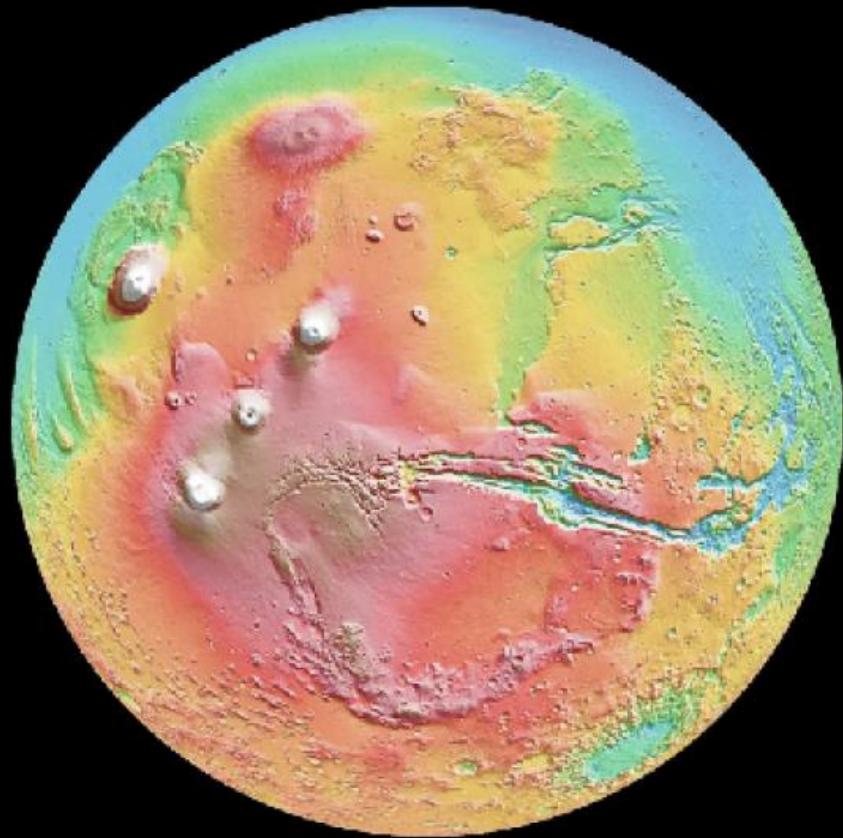
Mars

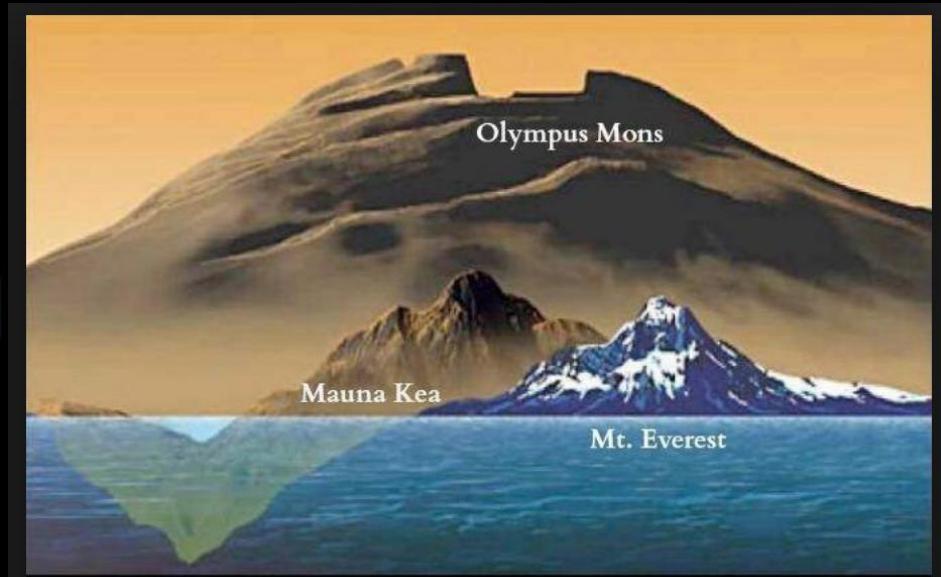
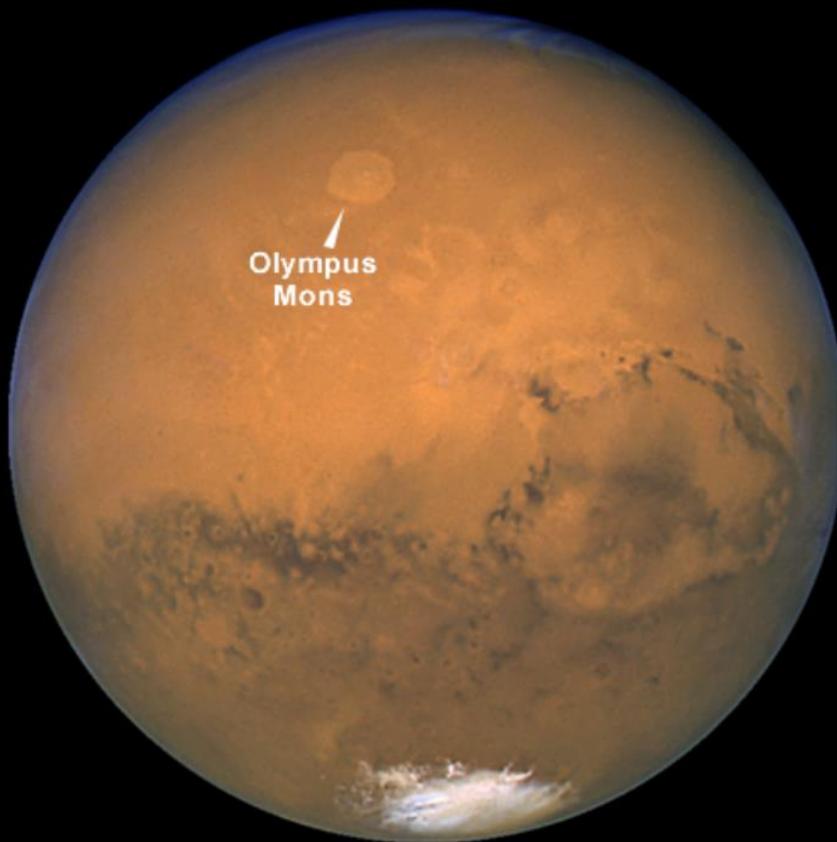


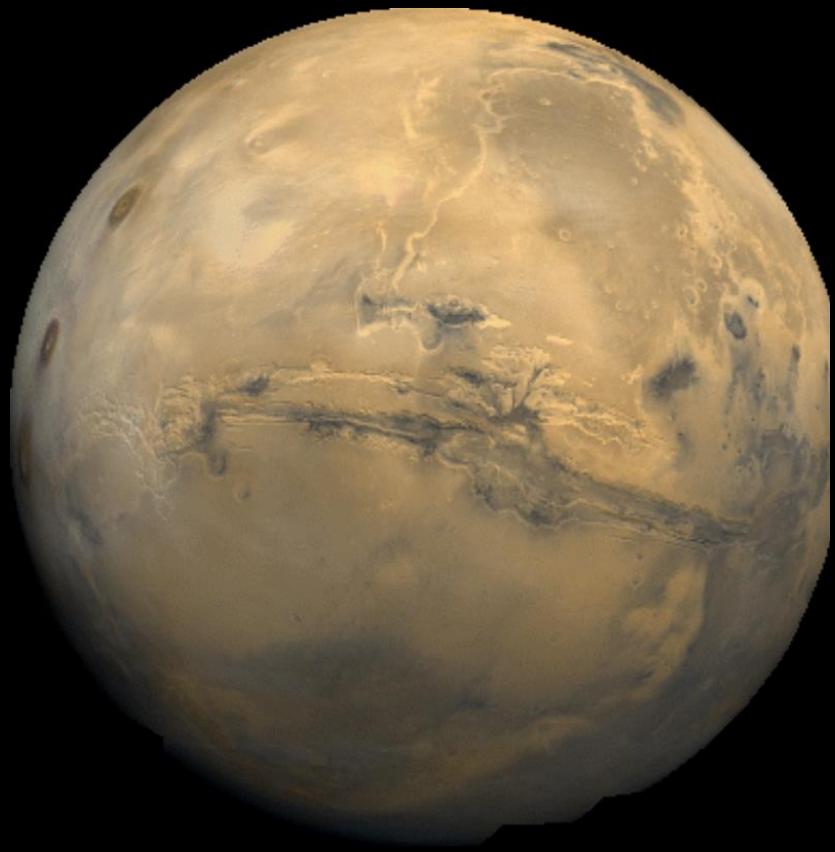
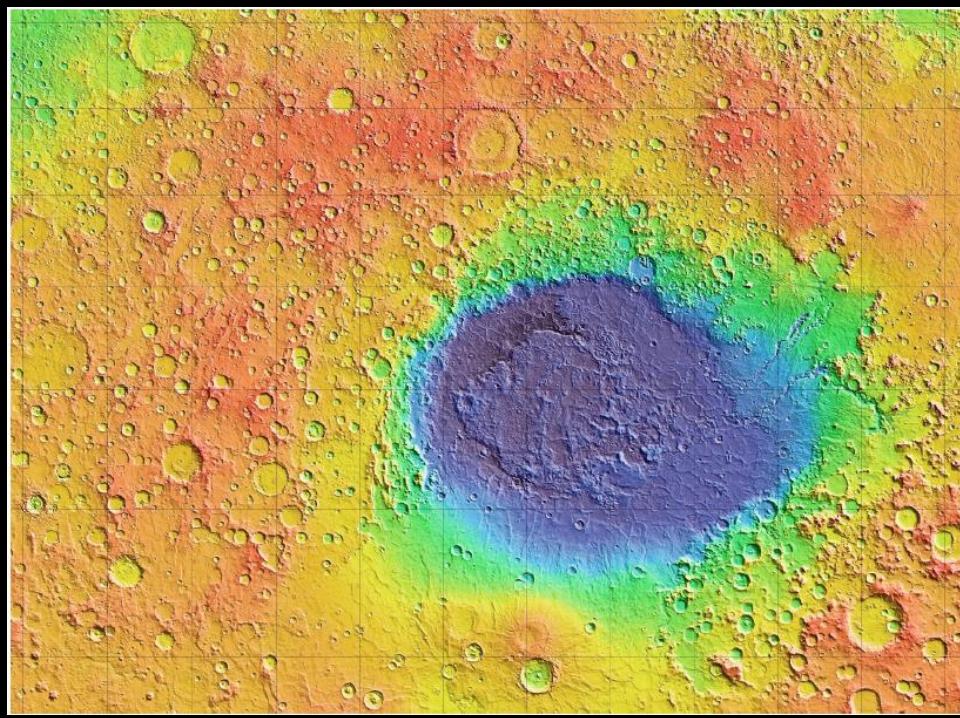
Mars rock from Antarctica!

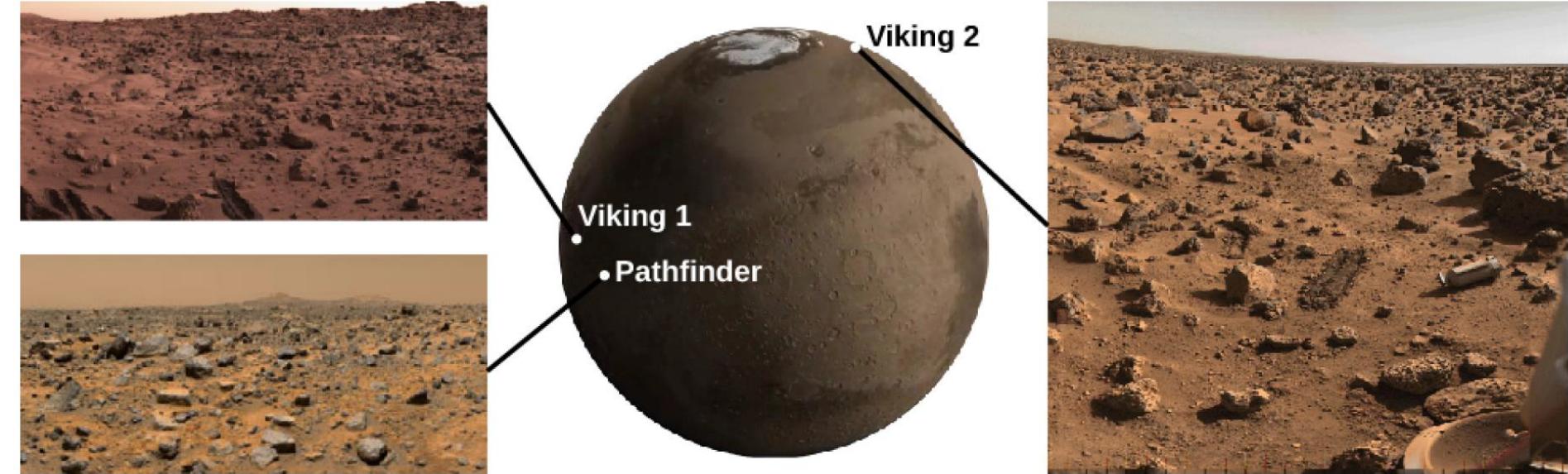




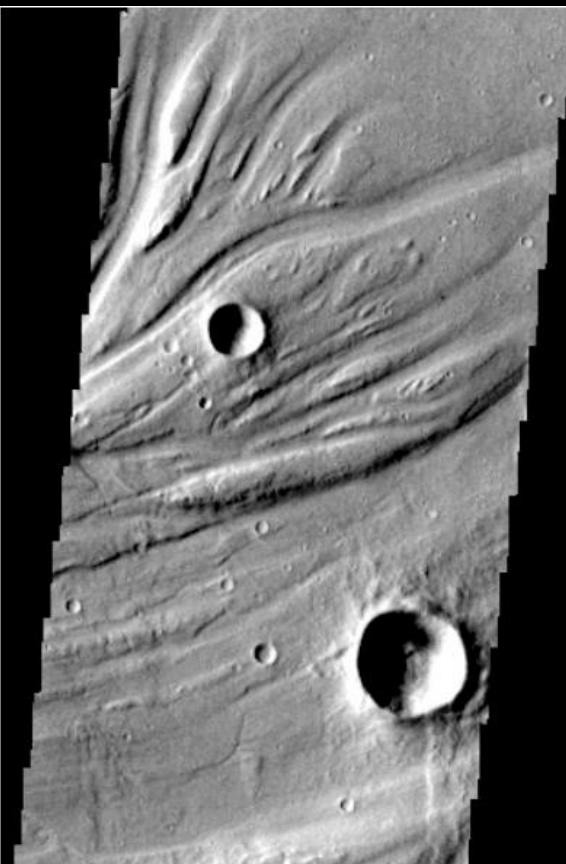


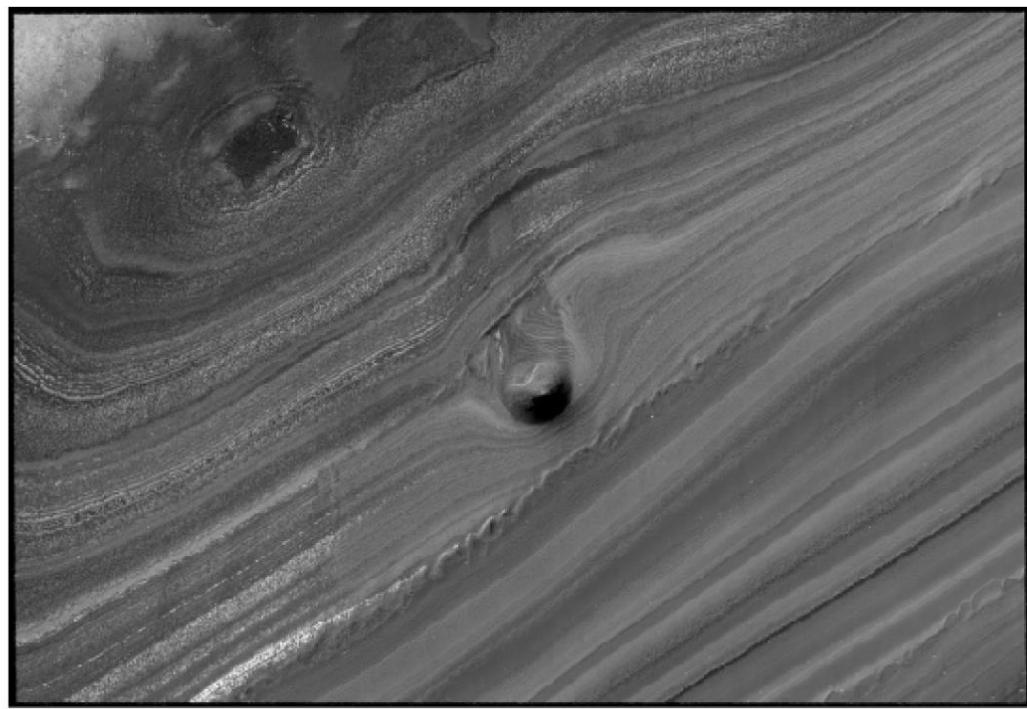
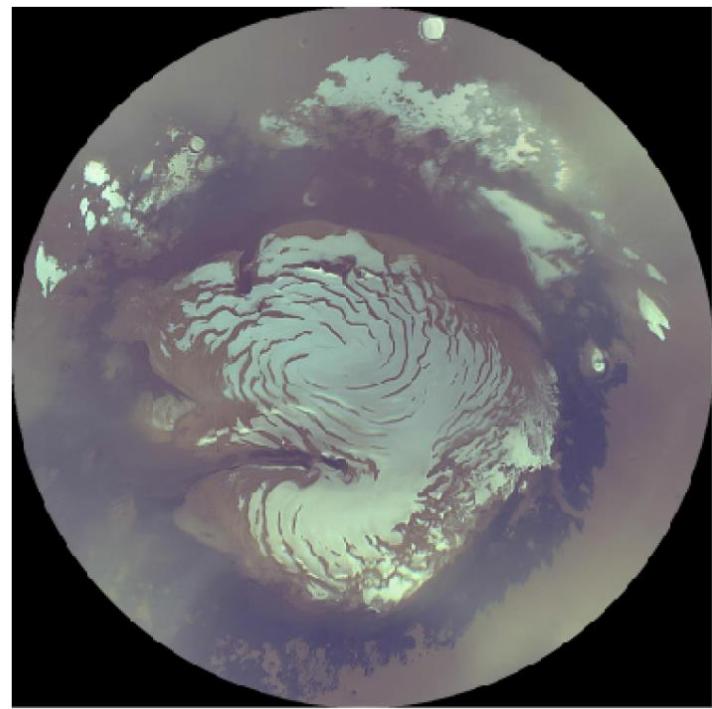


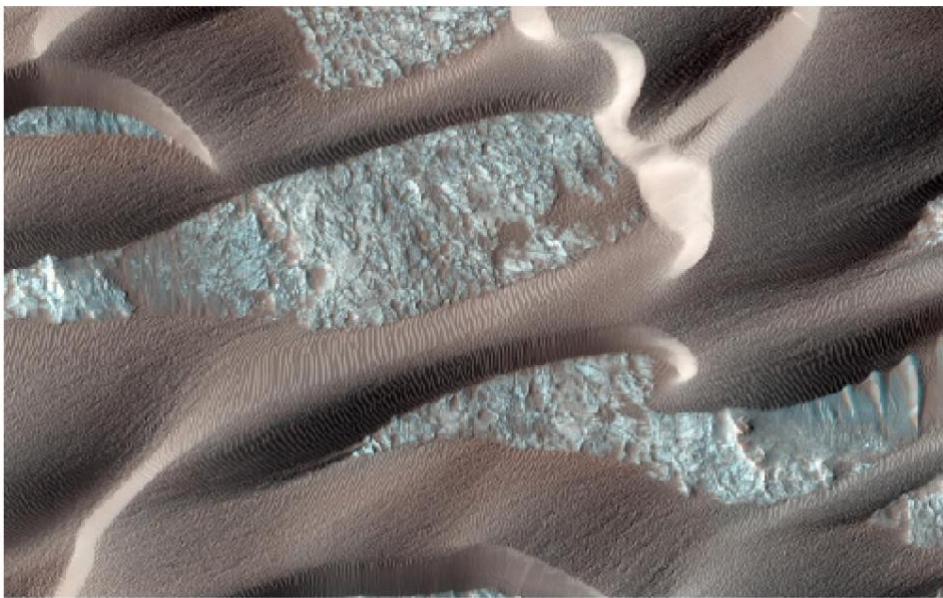


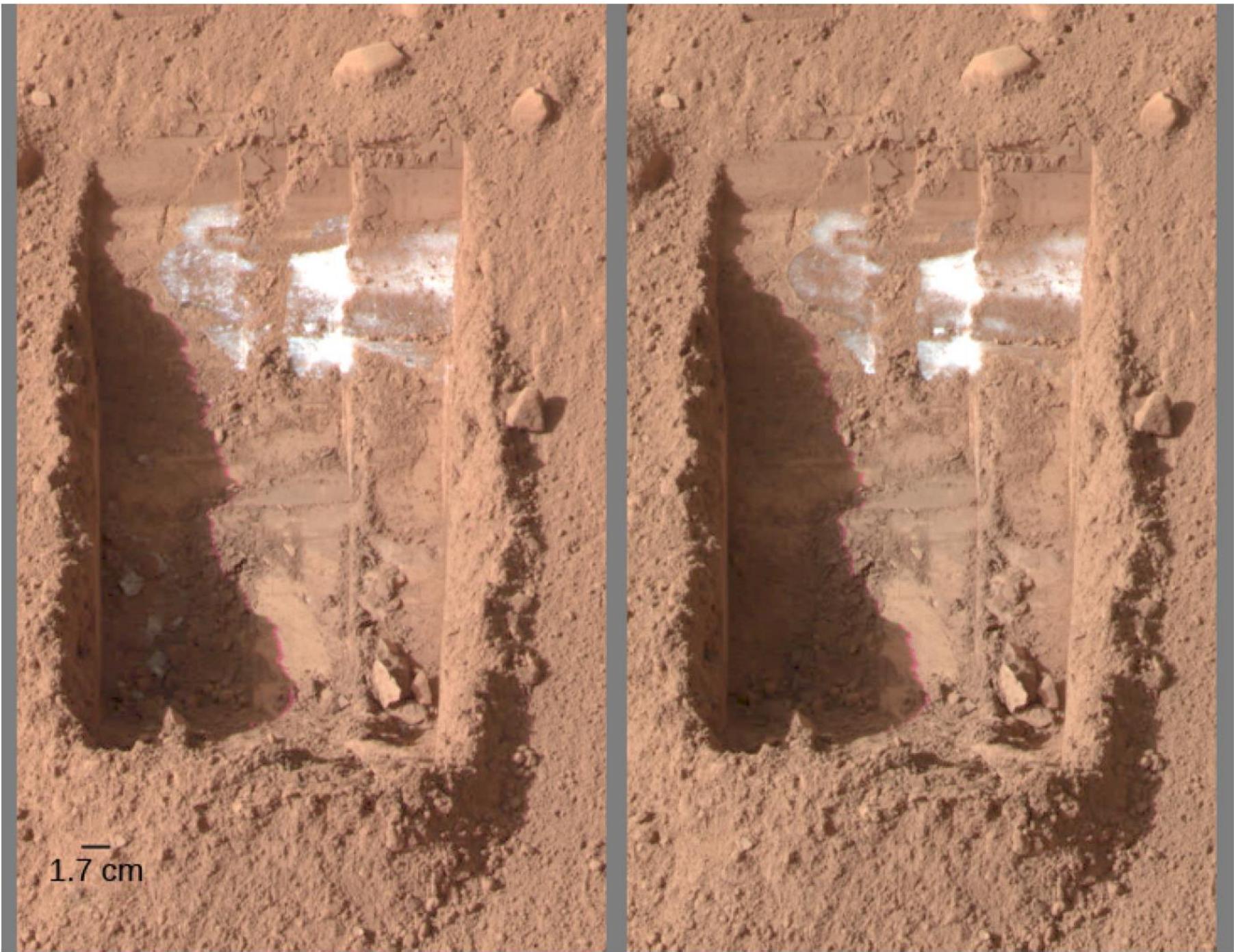


Water?

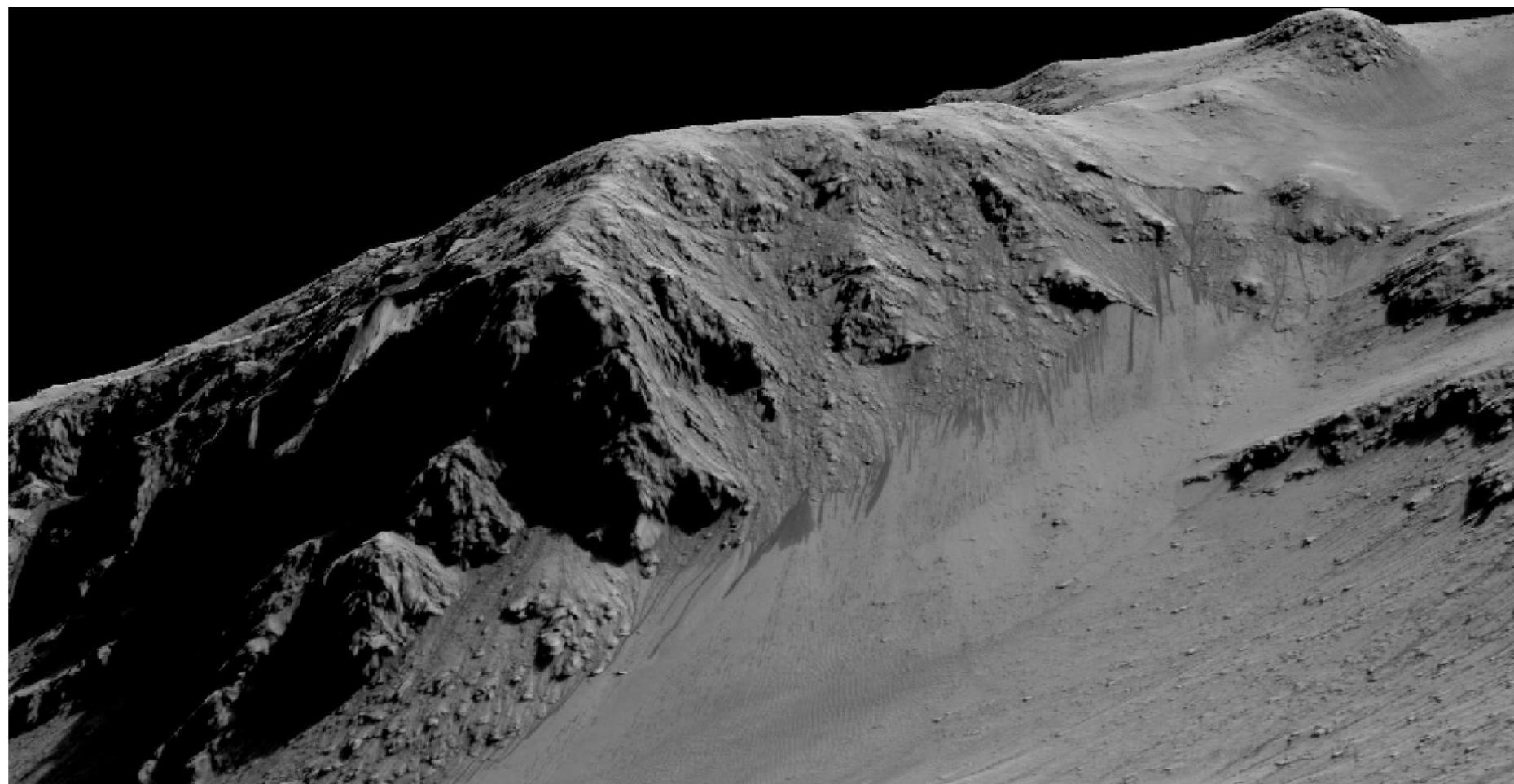






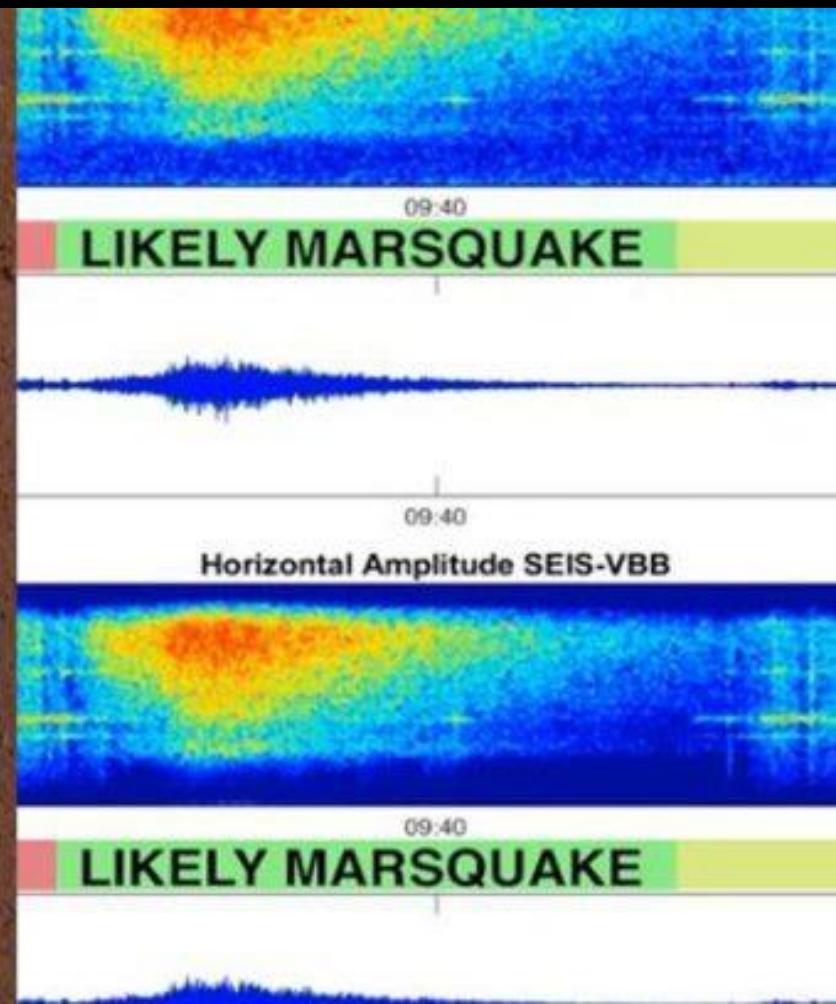


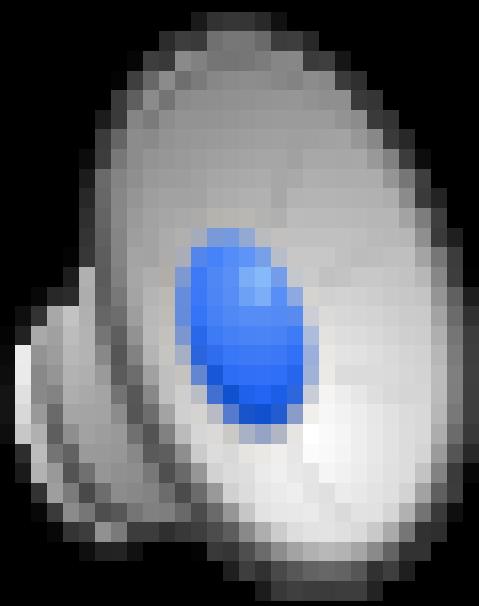
1.7 cm





Marsquakes!





KIM STANLEY
ROBINSON

Winner of the Nebula Award

Red Mars

'The ultimate in
future history'
Daily Mail

